The Synthesis and Structures of C₅H⁻ and C₅⁻⁻

Suresh Dua, John C. Sheldon and John H. Bowie

Department of Chemistry, The University of Adelaide, South Australia, 5005

The ion C_5H^- is formed in the gas phase by the process $-(C\equiv C)_2CH_2OMe \rightarrow C_5H^- + MeOH$; ab initio calculations indicate a ground state triplet structure with a near linear carbon skeleton, but with the terminal H at a marked angle to the carbon backbone.

Unusual organic molecules have been detected in circumstellar gas and dust envelopes which surround red giant stars.¹ These include C₃, C₅ and the hydrocarbons C_nH (n = 2-6), which have been identified around the bright carbon-rich star IRC + 10216.1-3 It is proposed that such neutrals are formed following photoionisation of acetylene, but the reaction sequences are not known.3 We are interested in synthesising anions corresponding to these neutrals and in investigating their structures and reactivities. We have already prepared C_3H^- in the gas phase by the collision induced process shown in eqn. (1). Ab initio

calculations indicate that C_3H^- has a ground state triplet diradical structure with allenic type geometry: the two unpaired electrons occupy the two lowest antibonding and orthogonal π orbitals which extend along the carbon chain, with most of the electron density being concentrated on the carbon which bears hydrogen.4

The linear C_5H^- ion has been synthesised by a similar and unambiguous route. Reaction of DC=C-C=CCH₂OMe with NH2⁻ in the chemical ionization source of a VG ZAB 2HF reverse-sector mass spectrometer yields an $(M - D)^{-}$ anion exclusively. A proportion of these parent anions have sufficient energy to undergo loss of MeOH to form an abundant C₅H⁻

species as the major decomposition product. The reaction is summarised in eqn. (2).

$$\label{eq:calculation} \begin{array}{l} \label{eq:calculation} \mbox{-}C \equiv \mbox{C} = \$$

When source-formed C_5H^- ions are fired through a collision cell (situated between the magnetic and electric sectors) containing He [at a pressure of 2×10^{-7} Torr (1 Torr is *ca*. 133 Pa)], the resultant negative-ion tandem mass spectrum (CID MS/MS) exhibits only loss of H[•], *i.e.* $C_5H^- \rightarrow C_5^{-} + H^{\cdot}$. Under precisely the same experimental conditions, collisional activation of C_5H^- also effects charge stripping to form C_5H^+ . Some of the resulting parent C_5H^+ ions have excess energy and undergo decomposition: the resultant spectrum is called a charge reversal mass spectrum [CR MS/MS (for descriptions of the technique, see refs. 5, 6)]. The CR MS/MS data for $C_5H^$ are [m/z (composition) relative abundance]: $61(C_5H^+)100$, $60(C_5^{+})76, 49(C_4H^+)27, 48(C_4^{+})28, 37(C_3H^+)27, 36(C_3^{+})33,$ $25(C_2H^+)1$, $24(C_2^{+*})3$, $13(CH^+)0.5$, $12(C^{+*})1$. This spectrum is quite characteristic, with C_5H^+ forming the major peak, together with less abundant peaks corresponding to the fragment cations C_n^{+} (n = 1-5) and C_nH^+ (n = 1-4). We have carried out *ab initio* calculations using GAUSSIAN

927 on C_5H^- and its decomposition product C_5^- : these data are

Table 1 Ab initio calculations for C_5H^- and C_5^{-a}

		$z \xrightarrow{y} C_1$	-C ₂ -C ₃ -C ₄ -C ₅ -H		
C ₅ H ⁻ Bond/Å C ₁ C ₂ C ₂ C ₃ C ₃ C ₄ C ₄ C ₅ C ₅ H	triplet 1.2854 1.3095 1.3003 1.2912 1.0615	singlet 1.2461 1.3308 1.2492 1.3236 1.0945	Angle/° $C_1-C_2-C_3(yz)^b$ $C_1-C_2-C_3(xz)^c$ $C_2-C_3-C_4(yz)$ $C_2-C_3-C_4(xz)$ $C_3-C_4-C_5(yz)$ $C_3-C_4-C_5(xz)$ $C_4-C_5-H(yz)$ $C_4-C_5-H(xz)$	<i>triplet</i> 180.0 180.0 177.4 180.0 173.2 178.7 150.2 185.7	singlet 180.0 180.0 178.5 180.0 175.7 179.5 110.6 190.7
HF/6-311G** Zero-point vibrational energy Total thermal energy (298.15 K) CISD/6-311G** C ₅ C ₁ C ₂ C ₃ C ₄ C ₅			triplet C ₁ symmetry -189.77870 ⁴ 0.02633 0.032080 -190.28989	singlet C ₁ symm -189.724 0.02855 0.03386 -190.275	etry 855 3 50 3615
Bond/Å C ₁ C ₂ C ₂ C ₃ C ₃ C ₄ C ₄ C ₅	1.2765 1.2950 1.2765 1.2950	D _{ooh} symme HF/6-311G [*] Zero-point v Total therm CISD/6-311	etry ** vibrational energy al energy (298.15 K) G**		15323 901 306 66782

a Ab initio optimization used the Berny procedure at HF6-311G**, GAUSSIAN 92.7 The correctness of the geometries was confirmed by the lack of negative analytical computed harmonic vibration frequencies. The post Hartree-Fock energies (CISD) were calculated at the Hartree-Fock geometries. ^b This is the angle in the plane of the page. ^c This is the angle in the plane perpendicular to the page. ^d All energies in hartrees.

collected in Table 1. There are triplet and singlet forms of C5Hwith the ground state corresponding to the triplet state [53 kJ mol⁻¹ lower in energy than the singlet (CISD level, zeropoint energy corrected and at 298.15 K)]. The energy necessary to remove an electron from triplet C_5H^- (i.e. the electron affinity of C₅H) is calculated as 229 kJ mol⁻¹. The atomic linkages are similar in both triplet and singlet states with the carbon atoms forming a near to linear chain, but with the hydrogen being at a marked angle to the carbon skeleton. The triplet state is characterised by small differences in carboncarbon distances along the chain [1.27-1.30 Å (these distances are similar to those reported⁸ for CH₂CCH₂ and OCCCO)]. In valence bond terms the geometry approximates to [C=C=C=C=C=CH]-: in MO terms, there are two unpaired electrons occupying the two lowest and mutually orthogonal antibonding π orbitals along the length of the carbon chain. Thus, the ground state structures of C_3H^- and C_5H^- are comparable. In contrast, the singlet state of C_5H^- is marked by alternation of carbon-carbon distances [1.25 and 1.33 Å (C=C distance in Me(C=C)₂H is 1.212 Å⁸)], suggesting that there is partial di-yne character, *i.e.* (C=C-C=C-CH)⁻. Finally, *ab* initio calculations indicate that the ground doublet state of the odd-electron species C_5^{-} is :C=C=Č=C=Č:

We thank the Australian Research Council for the financial support of this project.

Received, 24th February 1995; Com. 5/01146I

References

- 1 H. Olofsson, Molecular Abundances in the Envelopes around Evolved Stars, in Molecules in the Stellar Environment, Lecture Notes in Physics, ed. O. G. Jorgensen, Springer, Heidelberg, 1994, pp. 114–133, and references cited therein.
- 2 P. F. Bernath, K. H. Hinkle and J. J. Keady, Science, 1989, 244, 562; K. H. Hinkle, Infrared Spectroscopy and Molecules in Circumstellar Envelopes, in Molecules in the Stellar Environment, Lecture Notes in Physics, ed. O. G. Jorgensen, Springer, Heidelberg, 1994, pp. 99–114, and references cited therein.
- 3 A. Omont, Polyynes and Polycyclic Aromatic Molecules in C-rich Circumstellar Envelopes, in Molecules in the Stellar Environment, Lecture Notes in Physics, ed. O. G. Jorgensen, Springer, Heidelberg, 1994, pp. 135–148, and references cited therein.
- 4 S. Dua, J. H. Bowie and J. C. Sheldon, J. Chem. Soc., Perkin Trans. 2, 1994, 543.
- 5 J. H. Bowie and T. Blumenthal, J. Am. Chem. Soc., 1975, 97, 2959; J. E. Szulejko, J. H. Bowie, I. Howe and J. H. Beynon, Int. J. Mass Spectrom. Ion Phys., 1980, 13, 76.
- 6 R. Clair, J. L. Holmes, A. A. Mommers and P. C. Burgers, Org. Mass Spectrom., 1985, 20, 207; W. DeLange and N. M. M. Nibbering, Int. J. Mass Spectrom. Ion Processes, 1986, 68, 111.
- 7 M. J. Frisch, G. W. Trucks, M. Head-Gordon, P. M. W. Gill, M. W. Wong, J. B. Foresman, B. J. Johnson, H. B. Schlegel, M. A. Robb, E. S. Replogle, R. Gomperts, J. L. Andres, K. Raghavachari, J. S. Binkley, C. Gonzalez, R. L. Martin, D. J. Fox, D. J. Defrees, J. Baker, J. J. P. Stewart and J. A. Pople, GAUSSIAN 92, Revision G3, Gaussian Inc., Pittsburg, PA, 1992.
- 8 Tables of Interatomic Distances and Configuration in Molecules and Ions, The Chemical Society, London, 1958.