Inorg. Chem. **2005**, 44, 108−113



# **Infrared Spectrum of Hg(OH)<sub>2</sub>** in Solid Neon and Argon

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Received September 22, 2004

Mercury(II) hydroxide molecules have been prepared upon mercury arc lamp irradiation of Hg, H<sub>2</sub>, and  $O<sub>2</sub>$  mixtures in solid neon and argon. The strongest three infrared absorptions are identified through isotopic substitution  $(D_2,$ HD,  ${}^{18}O_2$ ,  ${}^{16}O{}^{18}O$ ) and comparison to frequencies from DFT calculations. The isolated Hg(OH)<sub>2</sub> molecule is stable and has a linear O−Hg–O linkage in a  $C_2$  structure with an 86° dihedral angle. However, in aqueous solution Hg<sup>2+</sup> and 2OH<sup>-</sup> may form an  $Hg(OH)_2$  intermediate, which eliminates water and precipitates solid HgO: The solid  $Hg(OH)<sub>2</sub>$  compound is not known.

#### **Introduction**

Mercury is an exotic element. Divalent mercury forms a number of binary compounds, and many of these are toxic.<sup>1-3</sup> The common  $HgCl<sub>2</sub>$  compound, calomel, is a volatile covalent molecular solid stabilized by "metallophilic" attraction, which gives a linear molecule in the vapor. $4-7$  The addition of OH<sup>-</sup> to Hg<sup>2+</sup> in aqueous solution precipitates the yellow solid HgO, which consists of zigzag  $-O-Hg O-Hg-$  chains with linear  $O-Hg-O$  units, and it is possible that  $Hg(OH)$ <sub>2</sub> is a transient intermediate in this precipitation reaction.<sup>1,8</sup> The Hg(OH)<sub>2</sub> molecule has been computed,<sup>9</sup> but there is no evidence for the gaseous molecule nor the solid compound<sup>1,2,8</sup> although  $Hg(OH)_2$  is thought to coprecipitate with Fe(OH)<sub>3</sub>.<sup>10</sup> The presence of mercury in industrial wastes is a significant environmental problem, and precipitation reactions provide a means of recovery. Furthermore, gas phase elemental mercury is rapidly oxidized to Hg(II) compounds in the arctic regions,<sup>11</sup> and  $Hg(OH)_2$  may play an intermediate role in this process.<sup>12</sup>

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The most closely related compound  $HgH<sub>2</sub>$  is reported to be an unstable solid, which decomposes at  $-125$  °C.<sup>13</sup> The  $HgH<sub>2</sub>$  molecule has been formed by 247 nm laser irradiation of Hg,  $H_2$  samples in excess argon, hydrogen, and nitro $gen<sup>14,15</sup>$  and by mercury arc lamp irradiation of Hg atoms in solid hydrogen.16 Recent work has shown that irradiation of mercury/water mixtures in excess argon gives the HHgOH molecule.17 Here we employ a mercury arc street lamp to form  $Hg(OH)_2$  molecules from mixtures of Hg, H<sub>2</sub>, and  $O_2$ in excess neon and argon and report the first experimental evidence for the  $Hg(OH)$ <sub>2</sub> molecule. In addition, we compare the infrared spectrum with vibrational frequencies calculated by density functional theory for the stable isolated  $Hg(OH)_{2}$ molecule.

#### **Experimental and Computational Methods**

The technique for reactions of metal atoms with hydrogen during condensation in excess hydrogen and neon has been described in detail previously.18,19 Mercury atoms were evaporated from the liquid at 50 °C in a stainless steel, valved system. Mercury atoms were thus codeposited with hydrogen  $(2 \text{ to } 8\%)$  and oxygen  $(0.2 -$ 2.0%) in excess neon, argon, or pure deuterium onto a 5 K CsI cryogenic window at  $2-4$  mmol/h (Ne, Ar for 1 h and pure  $D_2$  for

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**108** Inorganic Chemistry, Vol. 44, No. 1, 2005 10.1021/ic048673w CCC: \$30.25 © 2005 American Chemical Society Published on Web 11/18/2004

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Figure 1. Infrared spectra of neon matrix samples codeposited at 5 K with mercury vapor: (a) Ne and Hg with 2% H<sub>2</sub> and 0.2% O<sub>2</sub>, (b) after  $\lambda$ > 220 nm mercury arc irradiation for 25 min, (c) after annealing to 12 K, and (d) Ne and Hg with  $4\%$  H<sub>2</sub> after irradiation. W denotes water, and WC identifies water complex absorptions.

30 min). The  $D_2$ , HD (Cambridge Isotopic Laboratories),  ${}^{18}O_2$ , and  $16,18$ O<sub>2</sub> (Yeda) reagents were used in different experiments. FTIR spectra were recorded at  $0.5 \text{ cm}^{-1}$  resolution with  $0.1 \text{ cm}^{-1}$  accuracy on a Nicolet 750 using an HgCdTe detector. Matrix samples were annealed at different temperatures, and selected samples were subjected to broadband photolysis by a medium-pressure mercury arc street lamp (Philips, 175W) with globe removed.

Density functional theoretical calculations of mercury hydroxide are given for comparison. The Gaussian 98 program<sup>20</sup> was employed to calculate the structures and frequencies of expected product molecules using the B3LYP functional. The  $6-311++G(3df,3pd)$ basis set for O and H atoms and SDD pseudopotential for Hg were used. All the geometrical parameters were fully optimized, and the harmonic vibrational frequencies were obtained analytically at the optimized structures.

## **Results**

Infrared spectra and density functional calculations will be presented for the major products in the photochemical reaction of mercury with hydrogen and oxygen.

**Neon.** Figure 1 illustrates infrared spectra of the products formed by the reaction of mercury with hydrogen and oxygen in excess neon. The first sample with  $2\%$  H<sub>2</sub>, 0.2%  $O_2$ , and Hg vapor  $(0.1-0.2\%)$  codeposited at 5 K has no infrared absorption except for the usual trace impurities  $(H_2O, CO_2)$ , CO, CH<sub>4</sub>). Full mercury arc irradiation ( $\lambda > 220$  nm) for 25 min produces strong HgH<sub>2</sub> absorptions<sup>16</sup> at 1918.1 and 781.7  $cm^{-1}$  and new absorptions at 3702, 3642.3, 2118.2, 1612 (not shown), 928.8, and 644.2  $cm^{-1}$ . Note that the weak



**Figure 2.** Infrared spectra of neon and argon matrix samples codeposited with mercury vapor, hydrogen, and isotopic oxygen molecules, and irradiated at  $\lambda$  > 220 nm: (a) Ne, 2% H<sub>2</sub>, 0.2% <sup>16</sup>O<sub>2</sub>, (b) Ne, 3% H<sub>2</sub>, 0.4% <sup>16,18</sup>O<sub>2</sub>, (c) Ne, 2% H<sub>2</sub>, 0.3% <sup>18</sup>O<sub>2</sub> (c) Ne, 2% H<sub>2</sub>, 0.3% <sup>18</sup>O<sub>2</sub> (c) Ar, 10% H<sub>2</sub>, 1% (c) Ne, 2% H<sub>2</sub>, 0.3% <sup>18</sup>O<sub>2</sub>, (d) Ar, 8% H<sub>2</sub>, 0.8% <sup>16</sup>O<sub>2</sub>, (e) Ar, 10% H<sub>2</sub>, 1% <sup>16,18</sup>O<sub>2</sub>, and (f) 8% H<sub>2</sub>, 0.8% <sup>18</sup>O<sub>2</sub>. W denotes water.

 $3737.5$   $cm^{-1}$  water band does not grow on irradiation. Annealing to 12 K slightly increased the 3702 and  $1612 \text{ cm}^{-1}$ bands but decreased the other bands by 20-30%. Without oxygen the spectrum shows only the  $HgH<sub>2</sub>$  reaction product (Figure 1d). A similar experiment with Ne, Hg,  $O_2$ , and no H2 gave no observable product absorptions.

Similar investigations with  $^{18}O_2$  substitution shifted the new absorptions to 3695, 3631.7, 925.2 and 614.1 cm<sup>-1</sup>. The weak 2118.2 cm<sup>-1</sup> band was not shifted. A scrambled  $^{16,18}O_2$ sample gave doublet absorptions at  $3702$ ,  $3695$  cm<sup>-1</sup> and 3642, 3672 cm-<sup>1</sup> , a partially resolved triplet at 929, 927.1, 925 cm<sup>-1</sup>, and a triplet at 644.0, 632.3, 614.0 cm<sup>-1</sup>. Spectra showing the effect of oxygen isotopic substitution are compared in Figure 2, and the product absorptions are listed in Table 1.

Experiments with  $D_2$  and <sup>16</sup>O<sub>2</sub> gave strong HgD<sub>2</sub> absorptions at 1378 and 562  $\text{cm}^{-1}$  and new product absorptions at 2733, 2685.0, 1519.2, 711.5, and 619.9 cm<sup>-1</sup>. The new product bands again shifted with  ${}^{18}O_2$  as listed in Table 1, and  $16,18$ O<sub>2</sub> also gave doublet absorptions in the upper region and partially resolved triplet absorptions in the lower region, which are illustrated in Figure 3.

An investigation with  ${}^{16}O_2$  and HD was particularly productive: Strong absorptions were observed for HgHD at 1968, 1407, and 683  $cm^{-1}$ . New absorptions were observed at 3697, 3643.1, 2687.5, 934.3, 701.8, and 632.4 cm-1. The latter, lower frequency bands are also shown in Figure 3. Although no absorption was observed at  $2118 \text{ cm}^{-1}$ , a weak deuterium counterpart was observed at  $1520.6 \text{ cm}^{-1}$ .

**Argon.** Analogous experiments were done with higher reagent gas concentrations and Hg vapor in excess argon codeposited at 8 K. The samples were irradiated by the full mercury arc for 15 min, and HgH<sub>2</sub> absorptions<sup>14,21</sup> at 1895 and 773  $cm^{-1}$  and new product absorptions at 3629.4, 927.1, and  $637.3 \text{ cm}^{-1}$  were produced along with water and water dimer bands. Argon matrix bands were weaker and sharper

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Table 1. Infrared Absorptions (cm<sup>-1</sup>) Produced by Mercury Arc Irradiation of Solid Neon or Argon Containing Hydrogen, Oxygen, and Mercury

$H_2$ , <sup>16</sup> O <sub>2</sub>	$H_2$ , <sup>18</sup> O <sub>2</sub>	$H_2$ , 16,18 $O_2$	HD, $16O2$	$D_2$ , <sup>16</sup> O <sub>2</sub>	$D_2$ , <sup>18</sup> $O_2$	$D_2$ , 16,18 $O_2$	identity					
Neon Matrix												
3702	3695	3702, 3695	3697 br	2733	2717	2732, 2717	HOHX					
3642.3	3631.7	3642, 3632	3643.1, 2687.5	2684.0	2668.7	2685, 2668	Hg(OH) <sub>2</sub>					
2118.2	2118.2	2118.2	1520.6	1519.2	1519.2	1519.2	<b>HHgOH</b>					
1918	1918	1918	1968, 1407	1368	1368	1368	HgH <sub>2</sub>					
1612	1605	1612, 1605		1180	1171	1180, 1171	HOHX					
928.8	925.2	929, 927.1, 925	934.3	711.5	688.4	712, 704, 689	Hg(OH) <sub>2</sub>					
782	782	782	683	562	562	562	HgH <sub>2</sub>					
644.2	614.1	644.0, 632.3, 614.0	701.8, 632.4	619.9	598.2	620, 612, 599	Hg(OH) <sub>2</sub>					
551	524				512		[Hg(OH) <sub>2</sub> ]					
Argon Matrix												
3629.4	3618.6	3629.5, 3618.4	3629.7, 2677.9	2677.7			Hg(OH) <sub>2</sub>					
1896	1896	1986	1948, 1392	1363			HgH <sub>2</sub>					
927.1	923.7	927, 924.9, 924	932.1	704.5			Hg(OH) <sub>2</sub>					
773	773	773	675	555			HgH <sub>2</sub>					
637.3	607.8	637.3, 626.2, 607.8	698.9, 626.9				Hg(OH) <sub>2</sub>					

Table 2. Infrared Absorptions (cm<sup>-1</sup>) Produced by Mercury Arc Irradiation of Solid Deuterium Containing O<sub>2</sub> and Hg



than neon matrix counterparts. The samples were annealed to 12 K and photolyzed for another 20 min: The product absorptions doubled their intensity. The samples were annealed to 14 K and photolyzed again, and the product absorptions increased another 50%. The new product absorptions in solid argon are illustrated in Figure 2. Isotopic precursor molecules were reacted in a like manner; infrared spectra are compared in Figure 2, and band positions are listed in Table 1.

**Deuterium.** Mercury and  $O_2$  were subjected to irradiation in two solid deuterium samples, and the spectra revealed a different distribution of analogous product absorptions. Strong  $HgD_2$  bands at 1398 and 555 cm<sup>-1</sup> and  $D_2O$ absorptions at 2782, 2773.9, 2661.5, and 1178.3 cm<sup>-1</sup> and  $(D_2O)_2$  at 2611 cm<sup>-1</sup> were observed along with DO<sub>2</sub> at 2531.1, 1124.2, and 1024.6  $\text{cm}^{-1}$ .<sup>22-24</sup> Additional weak bands were observed at 2690.6, 2679.6, 1523.9, 709.4, 652, 617.9, and 569.1 cm-<sup>1</sup> . Continued irradiation substantially increased the HgD<sub>2</sub>, D<sub>2</sub>O, 2690.6, 1523.9, 652, and 569.1 cm<sup>-1</sup> absorptions and slightly increased the 2679.6, 709.4, and  $617.9 \text{ cm}^{-1}$  bands as shown in Figure 4. A similar irradiation of solid  $D_2$  containing  $O_2$  without Hg failed to produce  $D_2O$ .



Figure 3. Infrared spectra of neon matrix samples codeposited at 5 K with mercury vapor and irradiated at  $\lambda$  > 220 nm: (a) Ne and Hg with 2%  $H_2$  and 0.2%  $O_2$ , (b) Ne and Hg with 3% HD and 0.3%  $O_2$ , (c) Ne and Hg with 4%  $D_2$  and 0.4% <sup>16</sup>O<sub>2</sub>, (d) Ne and Hg with 3%  $D_2$  and 0.4% <sup>16,18</sup>O<sub>2</sub>, and (e) Ne and Hg with 2% and  $0.2\%$  <sup>18</sup>O<sub>2</sub>. WC indicates water complex.



**Figure 4.** Infrared spectra of deuterium matrix sample codeposited at 5 K with mercury vapor: (a)  $D_2$  and 0.1%  $O_2$ , (b) after  $\lambda > 220$  nm irradiation for 22 min, and (c) after  $\lambda$  > 220 nm irradiation for 20 additional min. X denotes common unknown photolysis product.

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Table 3. Vibrational Frequencies (cm<sup>-1</sup>) and Intensities (km/mol) of Hg(OH)<sub>2</sub> and HHgOH Calculated at the B3LYP/6-311++G(3df,3pd)/SDD Level

Hg(OH) <sup>a</sup>												
mode	HOHgOH	DOHgOD	HOHgOD	$H^{18}OHg^{18}OH$	$H^{16}OHg^{18}OH$	$D^{18}OHg^{18}OD$	$D^{16}OHg^{18}OD$					
$O-H$ str	3828.0(13)	2787.0(8)	3827.4(57)	3815.3(12)	3827.4(53)	2769.5(8)	2786.4(34)					
$O-H$ str	3826.8(101)	2785.8(65)	2786.4(36)	3814.2(98)	3814.8(59)	2768.4(61)	2768.8(36)					
$Hg-O-H$ bend	957.2(10)	706.2(99)	951.6(44)	953.6(9)	955.7(11)	698.0(6)	705.4(49)					
$Hg-O-H$ bend	946.1(81)	705.0(6)	705.7(56)	942.0(75)	943.7(76)	695.5(78)	697.0(46)					
$Hg-O str$	$606.9(112)^b$	588.7(67)	598.4(87)	578.2(104)	595.7(104)	566.0(74)	580.0(68)					
HHgOH <sup>c</sup>												
mode	<b>HHgOH</b>	DHgOD		HHgOD	<b>DHgOH</b>	HHg <sup>18</sup> OH	DHg <sup>18</sup> OD					
$O-H$ str	3846.3(42)	2800.1(29)		2800.1(29)	3846.2(42)	3833.6(41)	2782.6(27)					
$Hg-H$ str	2126.0(123)	1507.7(64)		2125.8(122)	1508.1(66)	2126.0(122)	1507.7(64)					
$Hg-O-H$ bend	871.8(40)		647.5(45)	688.9(30)	860.2(40)	868.1(38)	639.3(36)					
$H-Hg-O$ bend	584.7(5)		543.9(53)	583.2(7)	556.2(70)	583.8(4)	519.2(54)					
$H-Hg-O$ bend	579.2(11)	422.0(4)		557.8(7)	424.8(9)	578.8(12)	420.9(4)					
$Hg-O str$	556.2(74)	414.2(7)		525.1(2)	424.2(2)	527.5(67)	413.8(7)					

*a* Structure given in Figure 5 for <sup>1</sup>A ground state. *b* Four lower frequencies: 541.2(0), 167.0(5), 152.1(3), and 123.2(78) cm<sup>-1</sup> (km/mol). *c* Planar <sup>1</sup>A' ground state. H-Hg, 1.601 Å; Hg-O, 2.021 Å; O-H, 0.962 Å; <sup>∠</sup>HHgO, 176.7°; <sup>∠</sup>HgOH, 111.3°.



**Figure 5.** Structure of  $Hg(OH)_2$  calculated at the B3LYP/6-311++G-(3df,3pd)/SDD level. Bond lengths in angstroms and angles in degrees.

A like investigation of Hg and  ${}^{18}O_2$  in solid  $D_2$  gave the same HgD<sub>2</sub> bands,  $D_2$ <sup>18</sup>O, and  $D$ <sup>18</sup>O<sub>2</sub>: Additional weak bands are listed in Table 2. Continued irradiation behaved as described above.

**Calculations.** DFT calculations were performed on the two anticipated product species HHgOH and  $Hg(OH)_2$ , and the results are given in Table 3. Our structures are similar to those computed previously.<sup>9,17</sup> The <sup>1</sup>A  $Hg(OH)_2$  molecule has  $C_2$  symmetry, a linear O-Hg-O linkage, and a 85.8° dihedral angle as illustrated in Figure 5. The high energy OHgO molecule is calculated to be stable in a  ${}^{3}\Sigma_{g}^{-}$  state.<sup>25</sup>

### **Discussion**

The new absorptions will be assigned on the basis of observed isotopic shifts and splittings and comparison with isotopic frequencies computed by density functional theory.

**Hg(OH)2.** The major product requires three elements, Hg,  $H_2$ , and  $O_2$ , and the transient molecule  $Hg(OH)_2$  is identified from the mixed isotopic spectra shown in Figure 2. The 644.0, 632.3, 614.0  $cm^{-1}$  triplet in neon and the corresponding 637.3, 626.2, 607.8 cm<sup>-1</sup> triplet in argon using scrambled isotopic oxygen  $({}^{16}O_2 + {}^{16}O^{18}O + {}^{18}O_2)$  clearly demonstrate that two equivalent oxygen atoms are involved in this vibrational mode. The large  ${}^{16}O_2$  to  ${}^{18}O_2$  shift (16/18 ratio 1.0489 (neon) and 1.0485 (argon)) identifies this predominately antisymmetric HO-Hg-OH motion. The weaker  $928.8 \text{ cm}^{-1}$  absorption also gives a triplet band contour, and the small  ${}^{16}O_2$  to  ${}^{18}O_2$  shift and large H<sub>2</sub> to D<sub>2</sub> shift characterize an Hg-O-H bending mode. The associated band at 3642.3 cm<sup>-1</sup> shifts to 2684.0 cm<sup>-1</sup> with  $D_2$  in neon, and the corresponding argon matrix counterparts are 3629.4

and 2677.7 cm-<sup>1</sup> : These bands exhibit large H/D ratios  $(1.357 \text{ and } 1.355)$  that are characteristic of O-H stretching modes. Crucial information is obtained with the HD reagent, which gives new bands at  $3643.1$  and  $2687.5$  cm<sup>-1</sup> in neon and at 3629.7 and 2677.9  $cm^{-1}$  in argon in the upper region, and at 701.8 and 632.4  $cm^{-1}$  in neon and at 698.9 and 626.9  $cm^{-1}$  in argon in the lower region. Thus, with HD we observe two slightly different O-H and O-D stretching modes and  $Hg-O-H$  and  $Hg-O-D$  bending modes, and the  $Hg(OH)<sub>2</sub>$ molecule is thereby identified. There is precedent for dihydroxide molecules: the  $Be(OH)_2$  molecule is a major product of the laser-ablated Be atom and  $H_2O$  reaction,  $^{26}$  and the  $Ca(OH)_2$  molecule is formed by photolysis of Ca and  $H<sub>2</sub>O$  in excess argon.<sup>27</sup>

Assignment of the 3642.3, 928.8, and 644.2  $cm^{-1}$  absorptions to the  $Hg(OH)$ <sub>2</sub> molecule is confirmed by B3LYP harmonic frequency calculations, which predict the strongest three infrared absorptions at 3826.8, 946.1, and 606.9  $cm^{-1}$ . These predictions, 5.0% high, 1.9% high, 5.8% low, respectively, deviate more than previous results for first row transition metals,28 but the Hg-O stretching mode is difficult to model. The 18O shifts, however, are predicted accurately for these modes (calculated/observed:  $12.6/10.6$  cm<sup>-1</sup>, 3.6/ 3.6 cm<sup>-1</sup>, 28.7/28.1 cm<sup>-1</sup>), respectively.

**DHgOD.** The solid deuterium experiment was performed for the ultimate high concentration, and similar bands were observed at 2679.6, 709.4, and 617.9  $cm^{-1}$ , which are between the above neon and argon matrix observations for  $Hg(OD)$ . Three additional strong, sharp bands at 2690.6, 1523.9, and 569.1  $cm^{-1}$  and a weaker, broad band at 652  $cm^{-1}$  can be assigned to the related DHgOD molecule. First, the 1523.9, 652, and 569.1  $cm^{-1}$  bands are slightly higher than 1519.2, 646.7, and 565.6  $cm^{-1}$  argon matrix counterparts for the DHgOD product of the photochemical Hg +  $D_2O$ reaction where the  $2700 \text{ cm}^{-1}$  region was drowned in excess D<sub>2</sub>O.<sup>17</sup> Our B3LYP calculation predicts the O-D stretch 14.3  $cm^{-1}$  higher than that for  $Hg(OD)_2$ , and our new associated

<sup>(25)</sup> The OHgO  $(^3\Sigma_g^-)$  molecule has computed 1.894 Å Hg-O length and frequencies (intensities) 698 cm<sup>-1</sup> (15 km/mol) 625 (0) 184 (12 x frequencies (intensities) 698 cm<sup>-1</sup> (15 km/mol), 625 (0), 184 (12  $\times$ 2).

<sup>(26)</sup> Thompson, C. A.; Andrews, L. *J. Phys. Chem.* **1996**, *100*, 12214.

<sup>(27)</sup> Kauffman, J. W.; Hauge, R. H.; Margrave, J. L. *High Temp. Sci.* **1984**, *18*, 97.

<sup>(28)</sup> Bytheway, I.; Wong, M. W. *Chem. Phys. Lett.* **1998**, *282*, 219.

band at 2690.6 cm<sup>-1</sup> is 11.0 cm<sup>-1</sup> higher. The 16.3 cm<sup>-1</sup> observed <sup>18</sup>O shift is in good agreement with the 17.5 cm<sup>-1</sup> calculated value, which for a single O-D bond is necessarily less than the 19.7  $cm^{-1}$  <sup>18</sup>O shift observed for the sharp 2772.2 cm<sup>-1</sup> antisymmetric D<sub>2</sub>O stretching mode. The DHg-OD stretching mode observed here at  $569.1 \text{ cm}^{-1}$  shifts to  $544.7 \text{ cm}^{-1}$  with <sup>18</sup>O substitution, in excellent agreement with the calculated  $24.7 \text{ cm}^{-1}$  shift. Unfortunately, interaction with the  $D_2$  matrix broadened the 652 cm<sup>-1</sup> DHgOD bending mode.

Our B3LYP calculation predicts harmonic frequencies in good agreement with observed values for DHgOD as does a previous BP calculation.<sup>17</sup> The O-D stretch is predicted 4.0% high, which is in the range for this functional, but the Hg-D stretch and Hg-O stretch are predicted 1.0% and 4.4% too low. Our density functional calculations for mercury compounds are only approximate, but they support the recent argon matrix assignments.<sup>17</sup>

The strongest band, the Hg-H stretching mode, was observed for HHgOH at 2118.2 cm-<sup>1</sup> in solid neon, just 1.7  $cm^{-1}$  higher than the recent argon matrix observation,<sup>17</sup> and the DHgOD counterpart was observed at  $1519.2 \text{ cm}^{-1}$ . We note that with Hg, HD, and  $O_2$  in neon only the DHgOH isotopomer is formed. We provide additional information on the DHgOD transient species, namely the O-D stretching mode and <sup>18</sup>O shift for the Hg-O stretching mode, and its photochemical synthesis by a different route, which substantiates its recent identification.<sup>17</sup>

The HHgOH molecule may be compared with HXeOH formed in the 193 nm induced reaction of  $H_2O$  in pure xenon.29 The most interesting mode observed here for DHgOD in solid deuterium, the O-D stretching mode, was not observed for DXeOD, but the calculated O-D stretching frequency and bond length for these two species are nearly the same.

**Other Absorptions.** Two broad absorptions are produced at 3702 and  $1612 \text{ cm}^{-1}$  on irradiation in neon experiments. Their position near water absorptions invites our consideration of water, and their shift down from isolated water at  $3737.5$  cm<sup>-1</sup> and up from H<sub>2</sub>O at 1596.3 cm<sup>-1</sup> shows that hydrogen bonding is involved.<sup>30</sup> Hence, the 3702 and 1612  $cm^{-1}$  bands are due to a water complex  $HOH\cdots X$  where H2O and X are photolysis products. The isotopic shifts are in agreement with this assignment. As expected, the photochemical yield of water is highest in the solid  $D_2$  experiments. In addition,  $DO<sub>2</sub>$  radicals<sup>22,23</sup> are observed in the higher yield pure  $D_2$  investigations, which shows that  $D$  atoms are formed in the photochemical reactions with Hg atoms.

Finally, a weak 551 cm<sup>-1</sup> band shifts to 524 cm<sup>-1</sup> with  $18$ O and exhibits a 1.052  $16$ O/18O ratio, and the latter band has a 1.023 H/D ratio, which are comparable to the values observed for Hg(OH)<sub>2</sub>. This suggests that the 551 cm<sup>-1</sup> band is due to a Hg-O vibration, coupled with H, and the Hg-  $(OH)<sub>2</sub>$  dimer is a likely carrier for this band.



**Figure 6.** Relative energies computed at the B3LYP/6-311++G(3df,3pd)/ SDD level for important Hg,  $O_2$ , H<sub>2</sub> species.

**Reaction Mechanisms.** The mercury dihydroxide molecule is prepared here by the reaction of excited Hg with  $O<sub>2</sub>$ and H<sub>2</sub>, reaction 1, in the matrix cage. The excited  $Hg(^{3}P)$ state  $(+112.6 \text{ kcal/mol})^{31}$  is formed directly by mercury arc resonance radiation (254 nm) to activate this reaction. The stable <sup>1</sup>A Hg(OH)<sub>2</sub> molecule is calculated to be 54 kcal/mol lower in energy than  $Hg(^{2}S) + O_{2} + H_{2}$  at the B3LYP level.<br>In addition, we have evidence for H-O and HHgOH In addition, we have evidence for  $H_2O$  and  $HHgOH$ formation, which are likely products of reaction 2 with more dihydrogen. Molecular oxygen is not photodissociated at 254 nm,<sup>32</sup> and full arc irradiation of solid  $H_2$  containing  $O_2$  gives no photolysis products, but when Hg is added, HgH<sub>2</sub> is formed and  $O_2$  can react with both  $H_2$  and  $HgH_2$  to give H2O and HHgOH. Reaction 2 can, of course, give Hg and  $2H<sub>2</sub>O$  as well. Water and water dimer (Figures 2d and 4c) are also produced on full arc irradiation.

On the other hand, HXeOH is formed by a different mechanism in the 193 nm induced reaction of Xe and  $H_2O$ , which involves H and OH radicals.<sup>29</sup> We do not detect OH radicals<sup>33</sup> in this work, and OH is not a likely reagent in our 254 nm excited photochemical reactions.

$$
Hg(^{3}P) + O_{2} + H_{2} \rightarrow (OHgO)^{*} + H_{2} \rightarrow HOHgOH (1)
$$

$$
Hg(^3P) + O_2 + 2H_2 \rightarrow HHgOH + H_2O \tag{2}
$$

An interesting question is the mechanism of reaction 1. The reaction of Hg  $(^{3}P)$  and H<sub>2</sub> proceeds directly to HgH<sub>2</sub>, as shown in previous work $14-16$  and in Figure 1d. It is counterintuitive to suggest that reaction 1 proceeds through

<sup>(29)</sup> Petterson, M.; Khriachtchev, L.; Lundell, J.; Rasanen, M. *J. Am. Chem. Soc.* **1999**, *121*, 11904.

<sup>(30)</sup> Pimentel, G. C.; McClellan, A. L. *The Hydrogen Bond*; W. H. Freeman: San Francisco, 1960.

<sup>(31)</sup> Moore, C. E. *Atomic Energy Le*V*els*; Circular 467; National Bureau of Standards: Washington, DC, 1952.

<sup>(32)</sup> Huber, K. P.; Herzberg, G. *Constants of Diatomic Molecules*; Van Nostrand Reinhold: New York, 1979.

<sup>(33)</sup> Cheng, M. B.-M.; Lee, Y.-P.; Ogilvie, J. F. *Chem. Phys. Lett.* **1988**, *109*, 151.

 $HgH<sub>2</sub>$  although HgO is very weakly bound<sup>34</sup> and evidence for  $HgO<sub>2</sub>$  is tenuous.<sup>35</sup> Following the preparation of stable  ${}^{3}\Sigma_{g}^{-}$  states for the OZnO and OCdO molecules,<sup>36</sup> and the recent formation of stable  $Zn(OH)_2$  and  $Cd(OH)_2$  molecules in this laboratory, we have calculated a physically stable  ${}^{3}\Sigma_{g}^{-}$ state for OHgO, which is 72.5 kcal/mol higher in energy than Hg(<sup>1</sup>S) + O<sub>2</sub>, but this does not preclude OHgO ( ${}^{3}\Sigma_{g}$ <sup>-</sup>)<br>as a reactive intermediate in reaction 1. The relative as a reactive intermediate in reaction 1. The relative computed energies of important Hg,  $O_2$ , and H<sub>2</sub> species are given in Figure 6: Note that OHgO is 40 kcal/mol more stable than  $Hg(3P) + O_2$ . We suggest that OHgO produced<br>in the initial reaction then inserts into dihydrogen to form in the initial reaction then inserts into dihydrogen to form HOHgOH. Recall that no mercury oxides are observed when Hg and  $O_2$  are irradiated in excess neon. Hence,  $H_2$ chemically "traps" the OHgO subunit as the stable HOHgOH molecule. The strongest  $Hg(OH)_2$  absorption at 637.3 cm<sup>-1</sup> was also observed but not assigned in recent photochemical studies of Hg and  $H_2O$  in solid argon.<sup>17</sup> Therefore, the Hg-(OH)2 molecule is stable enough to result from the reaction of excited Hg and two water molecules.

Even though  $Hg(OH)_2$  is a very stable molecule, solid Hg- $(OH)_2$  does not precipitate from aqueous solution.<sup>1</sup> The HgO molecule is computed to be bound by about 4 kcal/mol, $34$ 

(37) Aurivillius, K. *Acta Chem. Scand.* **1964**, *18*, 1305.

and the total energy of HgO and  $H_2O$  molecules is approximately 57 kcal/mol above that for the  $Hg(OH)<sub>2</sub>$ molecule. Hence, the thermodynamic driving force for the precipitation of HgO from solution must arise mainly from the stability of the solid covalent chain zigzag HgO compound.37

## **Conclusions**

We have prepared the stable  $Hg(OH)_2$  molecule by mercury arc irradiation of Hg,  $H_2$ , and  $O_2$  mixtures in solid neon and argon. The strongest three infrared absorptions are identified through isotopic substitution  $(D_2, HD, {}^{18}O_2, {}^{16}O^{18}O)$ and comparison to frequencies from DFT calculations. The isolated  $Hg(OH)$ <sub>2</sub> molecule is stable and has a linear O-Hg-O linkage in a  $C_2$  structure with a 85.8° dihedral angle. A like irradiation of pure  $D_2$  doped with Hg and  $O_2$ formed  $Hg(OD)_2$  and DHgOD. The photochemical reaction is believed to proceed through a linear OHgO intermediate to the stable HOHgOH molecule. In aqueous solution,  $Hg^{2+}$ and  $2OH^-$  may form the  $Hg(OH)_2$  intermediate molecule, which eliminates  $H_2O$  and precipitates solid  $HgO$ ; hence, solid Hg(OH)<sub>2</sub> is not known.<sup>1,2,8</sup>

**Acknowledgment.** We gratefully acknowledge support from NSF Grant CHE 03-52487 and a preprint of ref 17 from T. M. Greene.

IC048673W

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<sup>(35)</sup> Butler, R.; Katz, S.; Snelson, A. *J. Phys. Chem.* **1979**, *83*, 2579. (36) Chertihin, G. V.; Andrews, L. *J. Chem. Phys.* **1997**, *106*, 3457.