

The Oxygen-Rich Carboxide Series: CO_n ($n = 3, 4, 5, 6, 7$, or 8)

Ben M. Elliott and Alexander I. Boldyrev*

Department of Chemistry and Biochemistry, Utah State University, Logan, Utah 84322-0300

Received: November 4, 2004; In Final Form: February 15, 2005

The lowest-energy isomers of the high-order carboxide series, CO_n ($n = 3–8$), have been elucidated via ab initio quantum chemical calculations. The structures of the lowest order of these (3 and 4) are in close agreement with previous calculations and with experimental data. The structures of the higher-order species are elucidated and correlated to the previous structures, showing similarities in structure and reactivity pathways. The reaction energies of the formation of all products are shown to be related. Exothermic pathways of formation often involve a C_{2v} form of CO_2 , which was shown to be metastable. The newly identified species could be intermediates in atmospheric reactions. The calculated vibrational frequencies and IR intensities may be used to identify these metastable species.

Introduction

Over the past few decades, there has been continual interest in high-order carboxides, CO_n ($n > 2$). Atmospheric studies in particular have been in the forefront of the research surrounding these species.^{1,2} The lowest order of these species, CO_3 , has been shown to be a key intermediate in reactions involving mono- and diatomic oxygen with low-order carboxides and organic molecules.³

High-order oxides of the group 12 elements would necessarily be highly energetic and, if they could be determined as metastable, would be very likely candidates for high-energy density materials (HEDMs).⁴ Wright et al. showed that high-order hydrogen oxides (HO_nH , with $n = 3–5$) could be a promising area of further research.⁵ The group of Plesničar^{6,7} and our own group⁸ characterized the HOOO^- species, which may be a precursor to oxygen dense compounds. Adamantides et al.⁹ theoretically predicted the covalently bound cyclic (D_{2d}) form of O_4 . Further theoretical effort occurred as this cyclic O_4 was ~5.3 eV higher in energy than two separate O_2 molecules.^{10–14} A D_{3h} form of O_4 analogous to SO_3 was found at a substantially higher energy, 6.5 eV.^{15,16} We have shown previously that charged, covalently bound tetra-oxygen O_4^{2-} ($C_{3v}, ^1\text{A}_1$) and O_4^{2+} ($D_{4h}, ^1\text{A}_{1g}$) and penta-oxygen O_5^{2-} species may be good candidates for oxygen-rich compounds.^{17–19}

Studies have been done on silicon oxide clusters,^{20,21} which show that this is a prospective area for further research. The C_{2v} cyclic isomer of CO_2 was found to be quite rigid and metastable with respect to the more common linear form.²² This gives the impression that structures based on a central carbon and bound oxygen rings may be feasible. The CO_3 molecule was much earlier trapped at temperatures of ~100 K.^{23,24} Recent theoretical studies were done on CO_3 ^{25,26} and CO_4 ,^{26–30} involving calculations at various levels of theory. Previous works predicted a C_{2v} isomer (Figure 1a) for CO_3 ,^{25,26} which was in close agreement with those earlier low-temperature experiments.

For the CO_4 molecule, two low-lying isomers (Figure 1b,c) were found as minima on the potential energy surface (PES).^{26–29} Cacace and co-workers recently reported the experimental detection of this CO_4 molecule.³⁰ They surmised, based upon their own and previous theoretical studies, that the lower-energy isomer, Figure 1b, would proceed to dissociation without a barrier, and that the D_{2d} structure, Figure 1c, would thus be the one observed. This work shows that this is a truly metastable molecule and is attractive for further research.

Herein, we describe the results of our ab initio investigation of the high-order carboxide series, CO_n ($n = 3–8$). Using high-level quantum chemical calculations, we report the geometries, stabilities, and harmonic frequencies of global minima and low-lying isomers of each of these species.

Computational Procedure

All calculations were performed using the Gaussian 03³¹ quantum chemical program package. The structures were visualized using the Molden 3.4³² molecular modeling program. The species' geometries were first optimized, and frequencies were calculated using the hybridized B3LYP^{33–35} method as the basic level of calculation with the split-valence 6-311+G*^{37–39} basis sets. Structures were then refined using the CCSD(T)^{39–41} method with the same basis sets. The appropriateness of the one-electron approximation methods was checked by using the CASSCF^{42,43} method with varying active-space sizes and the 6-311+G* basis to calculate geometries and harmonic frequencies.

Results and Discussion

CO_3 . The lowest-energy optimized structure is shown in Figure 1 (structure a), and it is a true local minimum. The optimized geometrical parameters and frequencies calculated at the various levels of theory are given in Table 1. Our results agree well with formerly published data.^{22–26}

The $C_{2v}, ^1\text{A}_1$ structure, at the CCSD(T) level of theory, has two C–O distances: the first is 1.18 Å, consistent with a C–O double bond, and the second is 1.33 Å, consistent with C–O single bonds. The O2–O3 distance, 1.65 Å, is somewhat

* To whom correspondence should be addressed. E-mail: Boldyrev@cc.usu.edu.

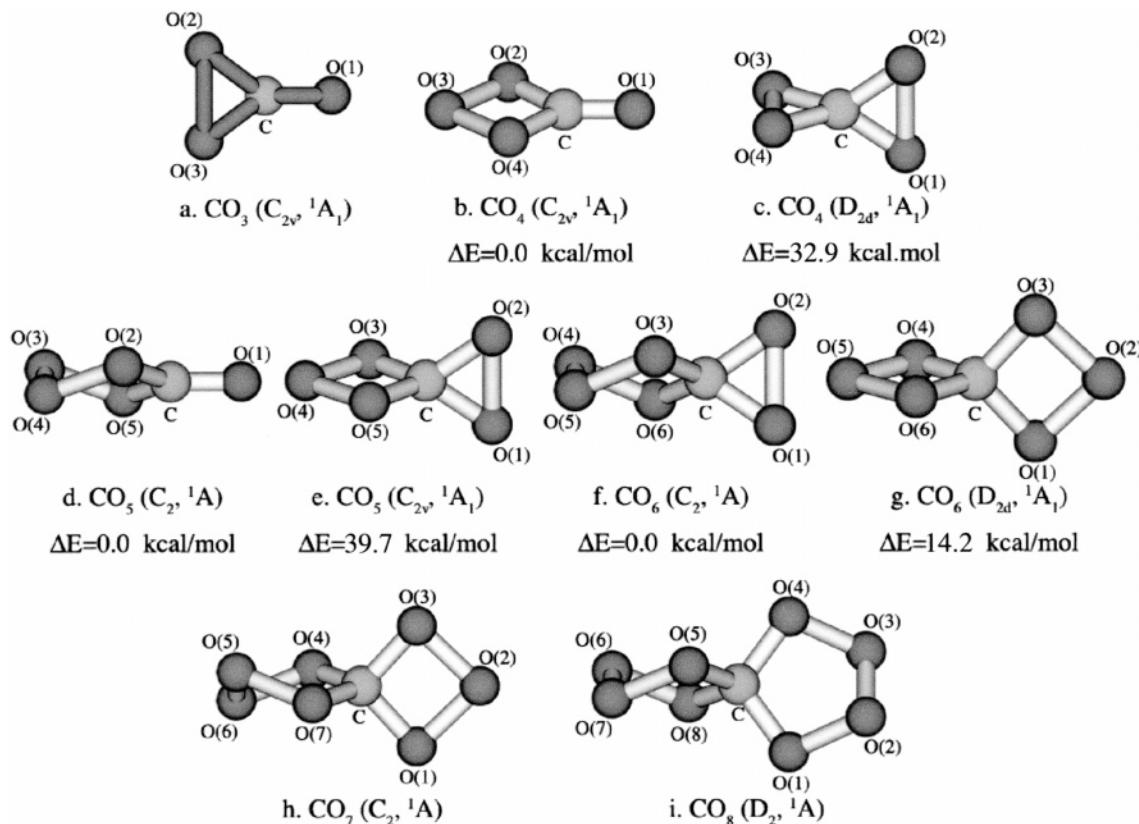


Figure 1. Structures of global minima and low-lying isomers of the CO_n ($n = 3, 4, 5, 6, 7$, or 8) series. The relative energies are given at the CCSD(T)/6-311+G* level of theory.

TABLE 1: Calculated Molecular Properties of CO_3

molecular parameter	B3LYP/6-311+G*	CASSCF(8,8)/6-311+G*	CCSD(T)/6-311+G*
CO_3 ($C_{2v}, ^1\text{A}_1$)			
E , au	-263.730 91	-262.514 11 {0.953} ^b	-263.149 65
$R(\text{C}-\text{O}_1)$, Å	1.17	1.15	1.18
$R(\text{C}-\text{O}_{2,3})$, Å	1.33	1.31	1.33
$R(\text{O}_2-\text{O}_3)$, Å	1.59	1.66	1.65
$\omega_1(\text{a}_1)$, cm^{-1}	2076 (631.1) ^a	2082	
$\omega_2(\text{a}_1)$, cm^{-1}	1122 (15.0)	1145	
$\omega_3(\text{a}_1)$, cm^{-1}	680 (12.9)	588	
$\omega_4(\text{b}_2)$, cm^{-1}	929 (74.6)	952	
$\omega_5(\text{b}_2)$, cm^{-1}	571 (6.4)	600	
$\omega_6(\text{b}_1)$, cm^{-1}	669 (35.2)	759	

^a The values in parentheses represent the intensities (km/mol) of IR active vibrational modes. ^b The values in braces indicate C_{HF} values in the CASSCF expansion.

overlong for an O–O single bond but may be tentatively attributed to the strain of the three corniced ring structure and the repulsion of the electrons on the electron-rich oxygens. The Hartree–Fock configuration is dominant in the CASSCF(8,8) wave function ($C_{\text{HF}} = 0.953$), indicating that our results based on the single-determinant methods (B3LYP and CCSD(T)) are reliable.

Although CO_3 is not stable against dissociation into CO_2 and O, at the B3LYP level of theory, the reaction of CO with O₂ exothermically yields CO_3 ($C_{2v}, ^1\text{A}_1$). At the highest level of theory (CCSD(T)), the value is again exothermic with a fair amount of released energy.

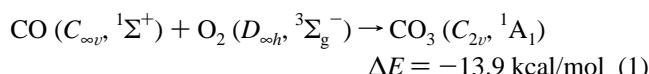


TABLE 2: Calculated Molecular Properties of CO_4

molecular parameter	B3LYP/6-311+G*	CASSCF(8,8)/6-311+G*	CCSD(T)/6-311+G*
CO_4 ($C_{2v}, ^1\text{A}_1$)			
E , au	-338.904 47	-337.305 97 {0.978} ^b	-338.155 77
$R(\text{C}-\text{O}_1)$, Å	1.17	1.15	1.18
$R(\text{C}-\text{O}_2)$, Å	1.37	1.34	1.37
$R(\text{O}_2-\text{O}_3)$, Å	1.46	1.50	1.48
$\angle(\text{O}_2-\text{C}-\text{O}_4)$, deg	95.5	98.0	96.5
$\omega_1(\text{a}_1)$, cm^{-1}	1996 (599.4) ^a	2073	
$\omega_2(\text{a}_1)$, cm^{-1}	1033 (42.4)	1189	
$\omega_3(\text{a}_1)$, cm^{-1}	912 (0.4)	1165	
$\omega_4(\text{a}_1)$, cm^{-1}	789 (13.5)	888	
$\omega_5(\text{b}_2)$, cm^{-1}	1116 (113.8)	1180	
$\omega_6(\text{b}_2)$, cm^{-1}	866 (1.0)	1161	
$\omega_7(\text{b}_2)$, cm^{-1}	513 (3.5)	455	
$\omega_8(\text{b}_1)$, cm^{-1}	722 (24.2)	791	
$\omega_9(\text{b}_1)$, cm^{-1}	298 (1.5)	306	
CO_4 ($D_{2d}, ^1\text{A}_1$)			
E , au	-338.841 36	-337.308 67 {0.896} ^b	-338.103 40
$R(\text{C}-\text{O}_1)$, Å	1.34	1.33	1.34
$R(\text{O}_1-\text{O}_2)$, Å	1.56	1.68	1.61
$\omega_1(\text{b}_2)$, cm^{-1}	1670 (487.8)	1718	
$\omega_2(\text{b}_2)$, cm^{-1}	753 (16.8)	458	
$\omega_3(\text{a}_1)$, cm^{-1}	1034 (0)	978	
$\omega_4(\text{a}_1)$, cm^{-1}	606 (0)	422	
$\omega_5(\text{e})$, cm^{-1}	1005 (49.3)	1186	
$\omega_7(\text{e})$, cm^{-1}	540 (12.5)	590	
$\omega_8(\text{b}_1)$, cm^{-1}	345 (0)	376	

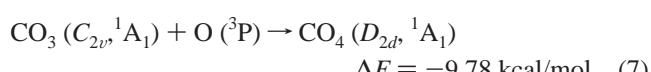
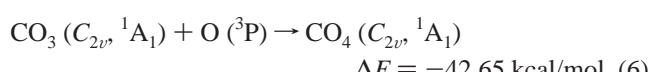
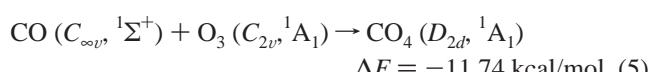
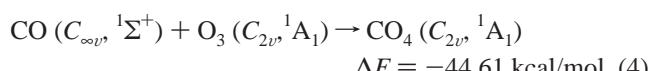
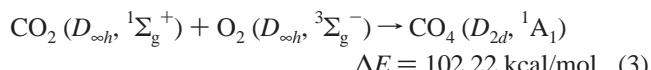
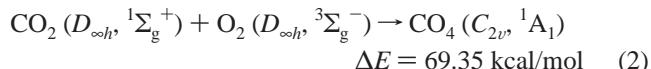
^a The values in parentheses represent the intensities (km/mol) of IR active vibrational modes. ^b The values in braces indicate C_{HF} values in the CASSCF expansion.

(All energy values for reaction pathways are given at the CCSD(T)/6-311+G* level of theory.)

CO₄. The lowest-energy optimized structures and relative energies are shown in Figure 1 (structures b and c). The optimized geometrical parameters and frequencies calculated at the various levels of theory are given in Table 2.

Concurrent with previous works,^{26–30} we find two isomers being local minima, with the C_{2v}, ¹A₁ isomer of CO₄ lying 30.8 kcal/mol below the D_{2d}, ¹A₁ isomer at the B3LYP level and 32.9 kcal/mol below it at the CCSD(T) level of theory. This energy difference could be attributed to the lack of π character to the bonding in the D_{2d} isomer as well as less ring strain in the four-membered ring of the C_{2v} isomer compared to the three-membered rings in the D_{2d} isomer. The Hartree–Fock configuration is again dominant in the CASSCF expansion for both structures (C_{HF}(C_{2v}) = 0.978, C_{HF}(D_{2d}) = 0.896).

These structures are both unstable with respect to dissociation into CO₂ and O₂ but are stable with respect to CO and O₃, as well as CO₃ and O.



CO₅. The lowest-energy optimized structures and relative energies are shown in Figure 1 (structures d and e). The optimized geometrical parameters and frequencies calculated at the various levels of theory are given in Table 3. No previous work could be found on these species according to the best of our knowledge.

Further addition of oxygen to the proposed series reveals two minima on the potential energy surface. The trend seen in the previous species is continued here. The lowest-energy isomer (Figure 1d) shows the characteristic doubly bonded oxygen opposite a five-membered ring. The bond lengths in the ring, at the CCSD(T) level, again correspond to typical C–O (1.38 Å) and O–O (1.46/1.42 Å) single bonds. The slightly increased energy difference ($\Delta E = 38.74$ kcal/mol, CCSD(T) level) between the two structures would probably correspond to the reduced ring strain of the five-membered ring and the energetics of the C–O double bond. Again, the Hartree–Fock configuration is dominant with C_{HF}(C₂) = 0.949 and C_{HF}(C_{2v}) = 0.901.

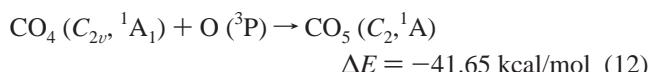
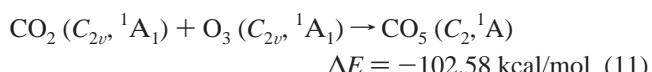
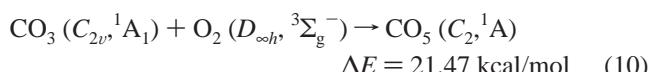
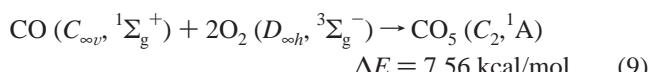
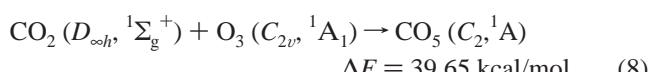
As with CO₄, we might suppose that the symmetry rules will allow the C_{2v} isomer to be much longer lived than the C₂ isomer. The formation energies of CO₅(C₂, ¹A) are endothermic via the typical pathways employing well-known molecules. However, the addition of a single oxygen atom to CO₄ and the

TABLE 3: Calculated Molecular Properties of CO₅

molecular parameter	B3LYP/ 6-311+G*	CASSCF(8,8)/ 6-311+G*	CCSD(T)/ 6-311+G*
CO ₅ (C ₂ , ¹ A)			
E, au	-414.072 33	-412.140 91 {0.949} ^b	-413.160 30
R(C–O ₁), Å	1.18	1.18	1.18
R(C–O ₂), Å	1.38	1.35	1.38
R(O ₂ –O ₄), Å	1.45	1.41	1.46
R(O ₃ –O ₄), Å	1.41	1.46	1.42
$\angle(\text{O}_2-\text{C}-\text{O}_5)$, deg	109.2	109.0	109.5
$\angle(\text{C}-\text{O}_2-\text{O}_4)$, deg	103.8	104.9	103.6
$\angle(\text{C}-\text{O}_2-\text{O}_4-\text{O}_3)$, deg	-38.3	-38.6	-38.3
$\omega_1(a)$, cm ⁻¹	1929 (479.7) ^a	1980	
$\omega_2(a)$, cm ⁻¹	939 (21.7)	1100	
$\omega_3(a)$, cm ⁻¹	886 (47.0)	1028	
$\omega_4(a)$, cm ⁻¹	762 (18.6)	971	
$\omega_5(a)$, cm ⁻¹	668 (6.7)	708	
$\omega_6(a)$, cm ⁻¹	389 (1.5)	407	
$\omega_7(b)$, cm ⁻¹	1092 (131.1)	1324	
$\omega_8(b)$, cm ⁻¹	867 (9.1)	810	
$\omega_9(b)$, cm ⁻¹	745 (18.3)	790	
$\omega_{10}(b)$, cm ⁻¹	696 (14.0)	765	
$\omega_{11}(b)$, cm ⁻¹	490 (3.9)	529	
$\omega_{12}(b)$, cm ⁻¹	198 (0.6)	178	
CO ₅ (C _{2v} , ¹ A)			
E, au	-414.000 90	-412.107 13 {0.901} ^b	-413.098 57
R(C–O _{1,2}), Å	1.35	1.34	1.35
R(C–O _{3,5}), Å	1.38	1.36	1.38
R(O _{3,5} –O ₄), Å	1.47	1.52	1.49
$\angle(\text{O}_1-\text{C}-\text{O}_2)$, deg	70.4	76.4	72.9
$\angle(\text{O}_3-\text{C}-\text{O}_5)$, deg	95.6	98.5	96.5
$\omega_1(a_1)$, cm ⁻¹	1495 (512.6)	1608	
$\omega_2(a_1)$, cm ⁻¹	994 (36.7)	1052	
$\omega_3(a_1)$, cm ⁻¹	930 (5.5)	752	
$\omega_4(a_1)$, cm ⁻¹	837 (19.5)	710	
$\omega_5(a_1)$, cm ⁻¹	590 (0.5)	463	
$\omega_6(b_1)$, cm ⁻¹	1147 (67.6)	1351	
$\omega_7(b_1)$, cm ⁻¹	908 (4.7)	734	
$\omega_8(b_1)$, cm ⁻¹	487 (6.4)	529	
$\omega_9(b_2)$, cm ⁻¹	958 (50.5)	1161	
$\omega_{10}(b_2)$, cm ⁻¹	556 (5.3)	628	
$\omega_{11}(b_2)$, cm ⁻¹	173 (0.4)	146	
$\omega_{12}(a_2)$, cm ⁻¹	284 (0)	350	

^a The values in parentheses represent the intensities (km/mol) of IR active vibrational modes. ^b The values in braces indicate C_{HF} values in the CASSCF expansion.

combination of the C_{2v} isomer of CO₂ with O₃ are both exothermic reactions.

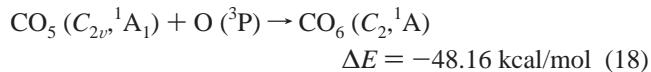
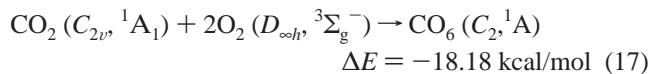
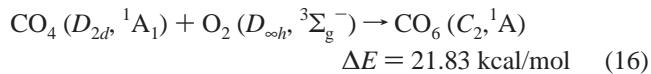
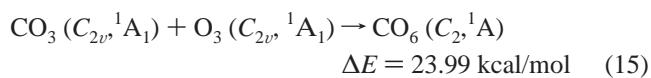
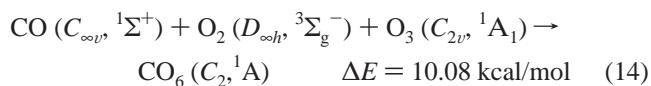
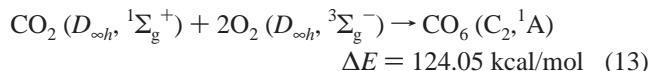


CO₆. The lowest-energy optimized structures and relative energies are shown in Figure 1 (structures f and g). The optimized geometrical parameters and frequencies calculated

at the various levels of theory are given in Table 4. The Hartree–Fock configurations in the CASSCF expansion (C_{HF}) are 0.930 and 0.963 for the C_2 and D_{2d} isomers, respectively.

In these structures, the trend of favoring the five-membered ring continues. Increased ring size beyond the five-membered ring, however, is not stable. A structure with a six-membered ring opposing a single oxygen–carbon double bond had one imaginary frequency and separated into CO_2 and O_2 upon following the frequency. This, along with the strain of the three-membered ring, makes the difference between the two isomers found to be only 14.2 kcal/mol (CCSD(T) values).

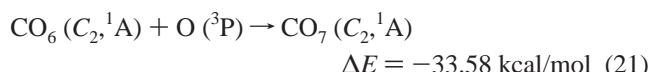
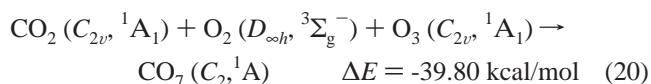
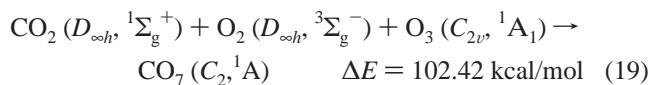
There is a similar situation here in the dissociation energies, as in CO_5 . The only exothermic channels involve CO_2 (C_{2v} , $^1\text{A}_1$) or O (^3P).



CO₇. The lowest-energy optimized structure is shown in Figure 1 (structure h). The optimized geometrical parameters and frequencies calculated at the various levels of theory are given in Table 5. The Hartree–Fock configuration is still dominant in the CASSCF wave function ($C_{HF} = 0.966$).

In this system, the only found minimum corresponds to an addition of O to the three-membered ring of the $\text{CO}_6(C_{2v})$ isomer, or to one of the four-membered rings of the D_{2d} isomer. The bond lengths are all familiar single O–O and C–O bonds.

The energetics follow the same pattern already seen.



CO₈. The lowest-energy optimized structure is shown in Figure 1 (structure i). The optimized geometrical parameters and frequencies calculated at the various levels of theory are given in Table 6. The CASSCF expansion gives a dominant Hartree–Fock configuration ($C_{HF} = 0.933$) once again.

Again, the only minimum found for this species was the addition of a single O to the previous $\text{CO}_7(C_2, ^1\text{A})$. The bond

TABLE 4: Calculated Molecular Properties of CO_6

molecular parameter	B3LYP/ 6-311+G*	CASSCF(8,8)/ 6-311+G*	CCSD(T)/ 6-311+G*
$\text{CO}_6(C_2, ^1\text{A})$			
E, au	-489.179 37	-486.854 99 {0.930} ^b	-488.113 47
$R(\text{C}-\text{O}_{1,2}), \text{\AA}$	1.35	1.34	1.35
$R(\text{C}-\text{O}_{3,6}), \text{\AA}$	1.39	1.37	1.40
$R(\text{O}_3-\text{O}_5, \text{O}_4-\text{O}_6), \text{\AA}$	1.45	1.36	1.46
$R(\text{O}_4-\text{O}_5), \text{\AA}$	1.40	1.47	1.41
$\angle(\text{O}_1-\text{C}-\text{O}_2), \text{deg}$	69.8	76.1	72.2
$\angle(\text{O}_3-\text{C}-\text{O}_6), \text{deg}$	108.4	106.7	108.5
$\angle(\text{C}-\text{O}_3-\text{O}_5), \text{deg}$	102.4	105.4	102.0
$\angle(\text{O}_3-\text{O}_5-\text{O}_4), \text{deg}$	99.5	98.7	98.9
$\angle(\text{C}-\text{O}_3-\text{O}_5-\text{O}_4), \text{deg}$	-42.5	-39.8	-43.7
$\omega_1(\text{a}), \text{cm}^{-1}$	1460 (376.7) ^a	1590	
$\omega_2(\text{a}), \text{cm}^{-1}$	948 (12.4)	1209	
$\omega_3(\text{a}), \text{cm}^{-1}$	883 (48.1)	991	
$\omega_4(\text{a}), \text{cm}^{-1}$	805 (2.5)	758	
$\omega_5(\text{a}), \text{cm}^{-1}$	709 (18.5)	656	
$\omega_6(\text{a}), \text{cm}^{-1}$	544 (5.1)	474	
$\omega_7(\text{a}), \text{cm}^{-1}$	472 (0.2)	446	
$\omega_8(\text{a}), \text{cm}^{-1}$	308 (0.5)	338	
$\omega_9(\text{b}), \text{cm}^{-1}$	1075 (82.8)	1295	
$\omega_{10}(\text{b}), \text{cm}^{-1}$	987 (43.3)	1204	
$\omega_{11}(\text{b}), \text{cm}^{-1}$	855 (9.4)	1161	
$\omega_{12}(\text{b}), \text{cm}^{-1}$	726 (1.1)	826	
$\omega_{13}(\text{b}), \text{cm}^{-1}$	575 (9.7)	641	
$\omega_{14}(\text{b}), \text{cm}^{-1}$	446 (3.0)	520	
$\omega_{15}(\text{b}), \text{cm}^{-1}$	136 (0.4)	135	
$\text{CO}_6(D_{2d}, ^1\text{A}_1)$			
E, au	-489.158 18	-486.788 63 {0.963} ^c	-488.090 91
$R(\text{C}-\text{O}_{1,3,4,6}), \text{\AA}$	1.39	1.36 ^c	1.39
$R(\text{O}_{1,3}-\text{O}_2, \text{O}_{4,6}-\text{O}_5), \text{\AA}$	1.47	1.44 ^c	1.48
$\angle(\text{O}_1-\text{C}-\text{O}_3), \text{deg}$	94.1	94.0 ^c	95.2
$\omega_1(\text{b}_2), \text{cm}^{-1}$	1268 (562.0)	1501 ^c	
$\omega_2(\text{b}_2), \text{cm}^{-1}$	944 (11.4)	1056 ^c	
$\omega_3(\text{b}_2), \text{cm}^{-1}$	880 (31.1)	1041 ^c	
$\omega_4(\text{e}), \text{cm}^{-1}$	1095 (67.5)	1305 ^c	
$\omega_5(\text{e}), \text{cm}^{-1}$	892 (3.1)	943 ^c	
$\omega_6(\text{e}), \text{cm}^{-1}$	500 (2.1)	585 ^c	
$\omega_7(\text{e}), \text{cm}^{-1}$	185 (0.3)	156 ^c	
$\omega_8(\text{a}_1), \text{cm}^{-1}$	1002 (0)	1131 ^c	
$\omega_9(\text{a}_1), \text{cm}^{-1}$	926 (0)	1042 ^c	
$\omega_{10}(\text{a}_1), \text{cm}^{-1}$	549 (0)	617 ^c	
$\omega_{11}(\text{b}_1), \text{cm}^{-1}$	214 (0)	329 ^c	

^a The values in parentheses represent the intensities (km/mol) of IR active vibrational modes. ^b The values in braces indicate C_{HF} values in the CASSCF expansion. ^c These values were calculated with the four active electrons and seven active molecular orbitals in the CASSCF expansion.

lengths are all very typical single bond lengths with no ring strain stretching. There are two exothermic pathways, one involving $\text{CO}_2(C_{2v}, ^1\text{A}_1)$ and $\text{O}_3(C_{2v}, ^1\text{A}_1)$ and the other CO_7 and $\text{O}(^3\text{P})$.

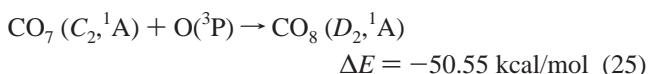
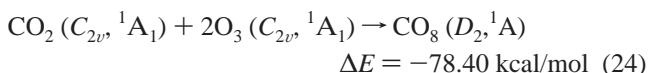
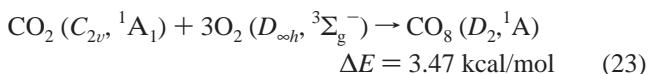
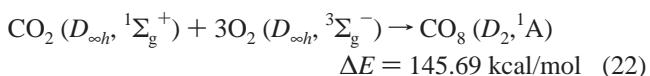


TABLE 5: Calculated Molecular Properties of CO₇

molecular parameter	B3LYP/ 6-311+G*	CASSCF(8,8)/ 6-311+G*	CCSD(T)/ 6-311+G*
CO ₇ (C ₂ , ¹ A ₁)			
E, au	-564.334 54	-561.623 09 {0.966} ^b	-563.105 13
R(C—O _{1,3}), Å	1.40	1.37	1.40
R(C—O _{4,7}), Å	1.40	1.37	1.40
R(O _{1,3} —O ₂), Å	1.47	1.39	1.49
R(O ₄ —O ₆ , O ₅ —O ₇), Å	1.45	1.40	1.46
R(O ₅ —O ₆), Å	1.40	1.44	1.41
∠(O ₁ —C—O ₃), deg	94.4	92.0	95.6
∠(O ₄ —C—O ₇), deg	108.8	106.6	107.1
∠(C—O ₄ —O ₆), deg	103.3	105.0	102.8
∠(O ₄ —O ₆ —O ₅), deg	99.2	98.6	98.7
∠(C—O ₇ —O ₅ —O ₆), deg	42.8	41.2	44.1
ω ₁ (a), cm ⁻¹	1221 (463.0) ^a	1464	
ω ₂ (a), cm ⁻¹	977 (1.2)	1190	
ω ₃ (a), cm ⁻¹	932 (0.3)	1134	
ω ₄ (a), cm ⁻¹	919 (11.5)	1060	
ω ₅ (a), cm ⁻¹	833 (30.0)	961	
ω ₆ (a), cm ⁻¹	708 (35.3)	798	
ω ₇ (a), cm ⁻¹	507 (3.6)	564	
ω ₈ (a), cm ⁻¹	455 (0.0)	463	
ω ₉ (a), cm ⁻¹	256 (0.2)	313	
ω ₁₀ (b), cm ⁻¹	1121 (52.3)	1312	
ω ₁₁ (b), cm ⁻¹	1037 (85.5)	1272	
ω ₁₂ (b), cm ⁻¹	909 (4.6)	1187	
ω ₁₃ (b), cm ⁻¹	844 (5.1)	1019	
ω ₁₄ (b), cm ⁻¹	711 (2.3)	811	
ω ₁₅ (b), cm ⁻¹	531 (4.9)	615	
ω ₁₆ (b), cm ⁻¹	465 (0.6)	541	
ω ₁₇ (b), cm ⁻¹	181 (0.6)	175	
ω ₁₈ (b), cm ⁻¹	100 (0.0)	99	

^a The values in parentheses represent the intensities (km/mol) of IR active vibrational modes. ^b The values in braces indicate C_{HF} values in the CASSCF expansion.

Alternative Structures

In the search for the global minimum, we calculated a large number of alternative structures. The CO_n species followed similar patterns until CO₇ was reached, at which point, the relatively low-lying alternative structures disappeared. We found that very few alternative structures for the higher-order species are minima on the potential energy surface. For example, any structure involving a six-membered ring was a saddle point with one imaginary frequency. Following this frequency caused all such species to fall apart into lesser species (e.g., CO₆ becomes CO₂ + 2O₂ and CO₈ becomes CO₄ + 2O₂). There were a few local minima on the PES. These species, however, proved to be so high on the PES as to be unfeasible. All other attempted species proved to be first-order or higher saddle points on the PES and following the imaginary frequencies led to either the found minima described above or to the dissociated van der Waals species.

Conclusions

We have found the lowest-energy covalently bound minima of the high-order carboxide series, CO_n (*n* = 3–8). The minima are very high-energy species lying far above the ground state of CO₂ (*D*_{∞h}, ¹Σ_g⁺) and O₂ (*D*_{∞h}, ³Σ_g⁻) and thus could be viable only as metastable species. There are few pathways for these species that would lead to them exothermically. However, we have shown that the use of CO₂ (C_{2v}, ¹A₁), which was shown to be a metastable higher-energy isomer of CO₂, for most of these species leads favorably to the desired products in the gas phase. The harmonic frequencies could be used for identifying the hyper-oxygen species, either under matrix

TABLE 6: Calculated Molecular Properties of CO₈

molecular parameter	B3LYP/ 6-311+G*	CASSCF(8,8)/ 6-311+G*	CCSD(T)/ 6-311+G*
CO ₈ (D ₂ , ¹ A ₁)			
E, au	-639.515 27	-636.486 29 {0.933} ^b	-638.123 84
R(C—O _{1,4,5,8}), Å	1.41	1.37	1.41
R(O ₁ —O ₂ , O ₅ —O ₇), Å	1.45	1.40	1.46
R(O ₂ —O ₃ , O ₆ —O ₇), Å	1.40	1.46	1.41
∠(O ₁ —C—O ₄), deg	107.3	106.8	107.8
∠(C—O ₁ —O ₂), deg	102.4	104.6	101.9
∠(O ₁ —O ₂ —O ₃), deg	99.3	98.2	98.7
∠(C—O ₁ —O ₂ —O ₃), deg	-43.4	-42.1	-44.6
ω ₁ (b ₂), cm ⁻¹	1181 (363.8) ^a	1417	
ω ₂ (b ₂), cm ⁻¹	940 (0.6)	1081	
ω ₃ (b ₂), cm ⁻¹	791 (18.3)	915	
ω ₄ (b ₂), cm ⁻¹	669 (85.9)	721	
ω ₅ (b ₂), cm ⁻¹	476 (1.1)	457	
ω ₆ (b ₁), cm ⁻¹	1074 (63.3)	1330	
ω ₇ (b ₁), cm ⁻¹	862 (7.6)	998	
ω ₈ (b ₁), cm ⁻¹	729 (3.7)	807	
ω ₉ (b ₁), cm ⁻¹	445 (0.0)	509	
ω ₁₀ (b ₁), cm ⁻¹	165 (0.5)	139	
ω ₁₁ (b ₃), cm ⁻¹	1018 (80.4)	1253	
ω ₁₂ (b ₃), cm ⁻¹	846 (1.2)	1003	
ω ₁₃ (b ₃), cm ⁻¹	711 (0.6)	802	
ω ₁₄ (b ₃), cm ⁻¹	556 (5.9)	638	
ω ₁₅ (b ₃), cm ⁻¹	83 (0.3)	57	
ω ₁₆ (a), cm ⁻¹	960 (0)	1109	
ω ₁₇ (a), cm ⁻¹	876 (0)	968	
ω ₁₈ (a), cm ⁻¹	719 (0)	740	
ω ₁₉ (a), cm ⁻¹	530 (0)	540	
ω ₂₀ (a), cm ⁻¹	428 (0)	463	
ω ₂₁ (a), cm ⁻¹	269 (0)	294	

^a The values in parentheses represent the intensities (km/mol) of IR active vibrational modes. ^b The values in braces indicate C_{HF} values in the CASSCF expansion.

isolation or in the gas phase under certain conditions. We hope that this research will lead to a broader view of the chemistries possible for high-order oxygen species.

Acknowledgment. This work was made possible by funds from the PRF administered by the ACS (ACS-PRF 38242-AC6).

References and Notes

- (a) Narcisi, R. S.; Bailey, A. D.; Della Lucca, L.; Sherman, C.; Thomas, D. M. *J. Atmos. Terr. Phys.* **1971**, 33, 1174. (b) Arnold, F.; Viggiano, A. A.; Ferguson, E. E. *Planet. Space Sci.* **1982**, 30, 1317. (c) Wisemberg, J.; Kockarts, G. *J. Geophys. Res.* **1980**, 85, 4642. (d) Thomas, L.; Bowman, M. R. *J. Atmos. Terr. Phys.* **1985**, 47, 547. (e) Viggiano, A. A.; Morris, R. A.; Paulson, J. F. *J. Chem. Phys.* **1989**, 91, 5855. (f) Dymek, M. K. *Adv. Space Explor.* **1980**, 8, 115. (g) Wayne, R. P. *Chemistry of Atmospheres*; Oxford University Press: Oxford, U.K., 2000.
- (a) Keating, G. M., Ed. *Adv. Space Res.* **1990**, 10, No. 5. (b) Fox, J. L. *Icarus* **1982**, 51, 248. (c) Hashimoto, G. L.; Abe, Y.; Sasaki, S. *Geophys. Res. Lett.* **1997**, 24, 289.
- (3) Parnis, J. M.; Hoover, L. E.; Fridgen, T. D.; Lafleur, R. D. *J. Phys. Chem.* **1993**, 97, 10708.
- (4) Shields, H.; Smith, A. L. S. *Appl. Phys.* **1978**, 16, 111. (b) Scott, W. B. *Aviat. Week Space Technol.* **1988**, 21, 19. (c) Klapötke, T. M.; Holl, G. *Green Chem.* **2001**, 3, G75.
- (5) Mckay, D. J.; Wright, J. S. *J. Am. Chem. Soc.* **1998**, 120, 1003.
- (6) Kraka, E.; Cremer, D.; Koller, J.; Plesničar, B. *J. Am. Chem. Soc.* **2002**, 124, 8462.
- (7) Plesničar, B.; Cerkovnik, J.; Tuttle, T.; Kraka, E.; Cremer, D. *J. Am. Chem. Soc.* **2002**, 124, 11260.
- (8) Elliott, B.; Alexandrova, A. N.; Boldyrev, A. I. *J. Phys. Chem. A* **2003**, 107, 1203.
- (9) Adamantides, V.; Neisius, D.; Verhaegen, G. *Chem. Phys.* **1980**, 48, 215.
- (10) Campargue, A.; Biennier, L.; Kachanov, A.; Jost, R.; Bussery-Honvault, B.; Veyret, V.; Churassy, S.; Bacis, R. *Chem. Phys. Lett.* **1988**, 288, 734.

- (11) Aquilanti, V.; Ascenzi, D.; Bartolomei, M.; Cappelletti, D.; Cavalli, S.; de Castro Vitores, M.; Pirani, F. *Phys. Rev. Lett.* **1999**, *82*, 69.
- (12) Peterka, D. S.; Ahmed, M.; Suits, A. G.; Wilson, K. J.; Korkin, A.; Nooijen, M.; Bartlett, R. J. *J. Chem. Phys.* **1999**, *110*, 6095.
- (13) Seidl, E. T.; Schaefer, H. F. *J. Chem. Phys.* **1988**, *88*, 7043; **1992**, *96*, 1176.
- (14) Dunn, K. M.; Scuseria, G. E.; Schaefer, H. F. *J. Chem. Phys.* **1990**, *92*, 6077.
- (15) Roeggen, I.; Nilssen, E. W. *Chem. Phys. Lett.* **1989**, *157*, 409.
- (16) Hottoka, M.; Pyykko, P. *Chem. Phys. Lett.* **1989**, *157*, 415.
- (17) Zhai, H.-J.; Yang, X.; Wang, X.-B.; Wang, L.-S.; Elliott, B.; Boldyrev, A. I. *J. Am. Chem. Soc.* **2002**, *124*, 6742.
- (18) Elliott, B. M.; Boldyrev, A. I. *Inorg. Chem.* **2004**, *43*, 4109.
- (19) Elliott, B. M.; Boldyrev, A. I. *J. Phys. Chem. A* **2005**, *109*, 236.
- (20) Chu, T. S.; Zhang, R. Q.; Cheung, H. F. *J. Phys. Chem. B* **2001**, *105*, 1705.
- (21) Lu, W. C.; Wang, C. Z.; Nguyen, V.; Schmidt, M. W.; Gordon, M. S.; Ho, K. M. *J. Phys. Chem. A* **2003**, *107*, 6936.
- (22) Xantheas, S. S.; Ruedenberg, K. *Int. J. Quantum Chem.* **1994**, *49*, 409 and references therein.
- (23) Weissberger, E.; Breckenridge, W. H.; Taube, H. *J. Chem. Phys.* **1967**, *47*, 1764.
- (24) Jacox, M. E.; Milligan, D. E. *J. Chem. Phys.* **1971**, *54*, 919.
- (25) Froese, R. D. J.; Goddard, J. D. *J. Phys. Chem.* **1993**, *97*, 7484.
- (26) Averyanov, A. S.; Khait, Y. G.; Puzanov, Y. V. *THEOCHEM* **1999**, *459*, 95.
- (27) Averyanov, A. S.; Khait, Y. G.; Puzanov, Y. V. *THEOCHEM* **1996**, *367*, 87.
- (28) Song, J.; Khait, Y. G.; Hoffmann, M. R. *J. Phys. Chem. A* **1999**, *103*, 521.
- (29) Korkin, A. A.; Schleyer, P. v. R.; Boyd, R. J. *Chem. Phys. Lett.* **1994**, *227*, 312.
- (30) Cacace, F.; Petris, G. d.; Rosi, M.; Troiani, A. *Angew. Chem., Int. Ed.* **2003**, *42*, 2985.
- (31) Frisch, M. J.; Trucks, G. W.; Schlegel, H. B.; Scuseria, G. E.; Robb, M. A.; Cheeseman, J. R.; Montgomery, J. A., Jr.; Vreven, T.; Kudin, K. N.; Burant, J. C.; Millam, J. M.; Iyengar, S. S.; Tomasi, J.; Barone, V.; Mennucci, B.; Cossi, M.; Scalmani, G.; Rega, N.; Petersson, G. A.; Nakatsuji, H.; Hada, M.; Ehara, M.; Toyota, K.; Fukuda, R.; Hasegawa, J.; Ishida, M.; Nakajima, T.; Honda, Y.; Kitao, O.; Nakai, H.; Klene, M.; Li, X.; Knox, J. E.; Hratchian, H. P.; Cross, J. B.; Adamo, C.; Jaramillo, J.; Gomperts, R.; Stratmann, R. E.; Yazyev, O.; Austin, A. J.; Cammi, R.; Pomelli, C.; Ochterski, J. W.; Ayala, P. Y.; Morokuma, K.; Voth, G. A.; Salvador, P.; Dannenberg, J. J.; Zakrzewski, V. G.; Dapprich, S.; Daniels, A. D.; Strain, M. C.; Farkas, O.; Malick, D. K.; Rabuck, A. D.; Raghavachari, K.; Foresman, J. B.; Ortiz, J. V.; Cui, Q.; Baboul, A. G.; Clifford, S.; Cioslowski, J.; Stefanov, B. B.; Liu, G.; Liashenko, A.; Piskorz, P.; Komaromi, I.; Martin, R. L.; Fox, D. J.; Keith, T.; Al-Laham, M. A.; Peng, C. Y.; Nanayakkara, A.; Challacombe, M.; Gill, P. M. W.; Johnson, B.; Chen, W.; Wong, M. W.; Gonzalez, C.; Pople, J. A. *Gaussian 03*, revision A.1; Gaussian, Inc.: Pittsburgh, PA, 2003.
- (32) Schaftenaar, G.; Noordik, J. H. *J. Comput.-Aided Mol. Des.* **2000**, *14*, 123.
- (33) Parr, R. G.; Yang, W. *Density-functional theory of atoms and molecules*; Oxford University Press: Oxford, U.K., 1989.
- (34) Becke, A. D. *J. Chem. Phys.* **1993**, *98*, 5648.
- (35) Perdew, J. P.; Chevary, J. A.; Vosko, S. H.; Jackson, K. A.; Pederson, M. R.; Singh, D. J.; Fiolhais, C. *Phys. Rev. B* **1992**, *46*, 6671.
- (36) Cizek, J. *Adv. Chem. Phys.* **1969**, *14*, 35.
- (37) Knowles, P. J.; Hampel, C.; Werner, H.-J. *J. Chem. Phys.* **1993**, *99*, 5219.
- (38) Frisch, M. J.; Pople, J. A.; Binkley, J. S. *J. Chem. Phys.* **1984**, *80*, 3265.
- (39) Raghavachari, K.; Trucks, G. W.; Pople, J. A.; Head-Gordon, M. *Chem. Phys. Lett.* **1989**, *157*, 479.
- (40) Clark, T.; Chandrasekhar, J.; Spitznagel, G. W.; Schleyer, P. v. R. *J. Comput. Chem.* **1983**, *4*, 294.
- (41) McLean, A. D.; Chandler, G. S. *J. Chem. Phys.* **1980**, *72*, 5639.
- (42) Bernardi, F.; Bottini, A.; McDougall, J. J. W.; Robb, M. A.; Schlegel, H. B. *Faraday Symp. Chem. Soc.* **1979**, *19*, 137.
- (43) Frisch, M. J.; Ragazos, I. N.; Robb, M. A.; Schlegel, H. B. *Chem. Phys. Lett.* **1992**, *189*, 524.