Reaction Mechanism of HCN⁺ + C₂H₄: A Theoretical Study

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The complex doublet potential energy surface for the ion-molecule reaction of HCN⁺ with C₂H₄ is investigated at the B3LYP/6-311G(d,p) and CCSD(T)/6-311++G(3df,2pd) (single-point) levels. The initial association between HCN⁺ and C₂H₄ forms three energy-rich addition intermediates, **1** (HCNCH₂CH₂⁺), **2** (HCcNCH₂CH₂⁺), and **3** (N-cCHCH₂CH₂⁺), which are predicted to undergo subsequent isomerization and decomposition steps. A total of nine kinds of dissociation products, including **P**₁ (HCN + C₂H₄⁺), **P**₂ (HCNCHCH₂⁺ + H), **P**₃ (NCCH₂ + CH₃⁺), **P**₄ (CN + C₂H₅⁺), **P**₅ (NCCHCH₂⁺ + H₂), **P**₆ (HNCCHCH₂⁺ + H), **P**₇ (c-CHCCH₂N⁺ + H₂), **P**₈ (c-NHCCH₂C⁺ + H₂), and **P**₉ (HNCCCH⁺ + H₂ + H), are obtained. Among the nine products, **P**₁ is the most abundant product. **P**₂ is the second feasible product but is much less competitive than **P**₁. **P**₃, **P**₄, **P**₅, and **P**₆ may have the lowest yields observed. Other products, **P**₇, **P**₈, and **P**₉, may become feasible at high temperature. Because the intermediates and transition states involved in the most favorable pathway all lie below the reactant, the HCN⁺ + C₂H₄ reaction is expected to be rapid, which is confirmed by experiment. The present calculation results may provide a useful guide for understanding the mechanism of HCN⁺ toward other unsaturated hydrocarbons.

1. Introduction

Titan, the largest satellite of Saturn, has attracted considerable interest because its atmosphere is so dense and because it is one of the places where the most complex atmospheric organic chemistry takes place in the solar system. A great deal of effort has been expended in describing the structure and the composition of Titan's atmosphere.¹⁻⁹ It has been established that Titan's dense and complex atmosphere is essentially composed of nitrogen and methane but also contains traces of hydrogen, ethane, ethylene, and acetylene.10-12 Primary ionization begins in Titan's ionosphere both by photoionization from solar radiation and by impact ionization from magnetospheric electrons.¹³⁻¹⁶ The reactions that occur via the primary ions N⁺ and N₂⁺ reacting with CH₄ initiate the ion-process, which yields the product ions CH_2^+ , CH_3^+ , CH_4^+ , and HCN⁺. The ions then react further with hydrocarbons that are present in abundance in Titan's atmosphere such as methane (CH₄), acetylene (C_2H_2), ethylene (C_2H_4), ethane (C_2H_6), and so on. In this way, a complex matrix of reactions is quickly established with a wide range of products. Such reactions are generally very fast and may be very effective in depleting old molecules or ions and in synthesizing new molecules or ions.

The ion-neutral chemistry in Titan's atmosphere is complex and has attracted the attention of chemists and astronomers. Up to now, a large number of experimental^{17–27} and theoretical^{28–32} studies have been performed to simulate the chemistry occurring in Titan's atmosphere. As one of the important ions in Titan's ionosphere, HCN⁺ has drawn considerable attention, and many ion-molecule reactions relevant to HCN⁺ have been investigated both experimentally and theoretically. In 2004, Anicich et al. performed experimental studies of HCN⁺ with CH₄, C₂H₂, and C₂H₄ using the flowing after low-selected ion flow tube (FA-SIFT) at room temperature.³³ Recently, we investigated the reactions of HCN⁺ with CH₄³⁴ and C₂H₂³⁵ using theoretical methods. As a part of our continuing interest in the gas-phase chemistry of HCN⁺ and because there has been no theoretical study on the HCN⁺ + C_2H_4 reaction up to now, we set out to investigate the reaction mechanism of HCN^+ with C_2H_4 in the present article.

2. Computational Methods

All calculations are performed with the Gaussian 98 program package.36 The geometries of all of the reactants, products, intermediates, and transition states are optimized using the popular density functional theory B3LYP37 method (the Becke's threeparameter hybrid functional with the nonlocal correlation functional of Lee-Yang-Parr) in conjunction with the d,p-polarized 6-311G (d,p) basis set.³⁸⁻⁴⁰ Frequency calculations are performed at the same level to check to see if the obtained species is an isomer or a transition state with an imaginary frequency of 0 and 1, respectively. To get more reliable energetic data, single-point energy calculations are carried out at the CCSD(T)/ $6-311++G(3df,2pd)^{41}$ level (coupled cluster approach with single and double substitutions including a perturbative estimate of connected triple substitutions) using the B3LYP/6-311G(d,p) optimized geometries. Finally, the intrinsic reaction coordinate42 calculations are performed at the B3LYP/6-311G(d,p) level to confirm that the transition states connect designated intermediates.

Our choice of B3LYP is motivated by its good performance for the geometry optimizations and prediction of vibrational frequencies for both isomers and transition states based on much previous literature.^{43–45} Moreover, B3LYP is able to suppress effectively the problem of spin contaminants. On the other hand, for the current reaction, the 6-311G(d,p) basis set is a balanced choice in consideration of computational efficiency and accuracy.

3. Results and Discussion

The optimized structures of the reactant and products are shown in Figure 1, and the optimized structures of intermediates and transition states are shown in Figures 2 and 3, respectively. The symbol **TSm/n** is used to denote the transition state connecting intermediates **m** and **n**. The energy of the reactant (${}^{2}\text{HCN}^{+} + {}^{1}\text{C}_{2}\text{H}_{4}$) is set to zero for reference. Unless otherwise indicated, the



Figure 1. Optimized structures of the reactant and products. Distances are given in angstroms and angles are given in degrees.

relative energies mentioned hereafter refer to the CCSD(T)/6-311++G(3df,2pd)//B3LYP/6-311G(d,p)+ZPVE (zero-point vibrational energy) level. The energetics of the reactant and products are listed in Table 1, whereas those of the intermediates and transition states are summarized in Table 2. By means of the interrelations among the reactant, intermediates, transition states, products, and corresponding relative energies, the schematic profile of the potential energy surface (PES) for the HCN⁺ + C₂H₄ reaction is depicted in Figure 4. **3.1. Initial Association.** The reactant HCN⁺ has $C_{\infty\nu}$ symmetry and a ² Σ electronic state. The spin densities on carbon and nitrogen are 0.419427e and 0.594288e respectively, as confirmed by theoretical calculations at the B3LYP/6-311G(d,p) level. So, both carbon and nitrogen can be viewed as the active site of HCN⁺. Via the attack of the nitrogen atom of HCN⁺ at one carbon atom of C₂H₄, intermediate **1** (HCNCH₂CH₂⁺) can be formed. Alternatively, the nitrogen atom can attack the C=C double bond of C₂H₄, and a three-membered ring intermediate **2** (HC-cN CH₂-





Figure 2. Part 2 of 2. Optimized structures of the intermediates. Distances are given in angstroms and angles are given in degrees.

 CH_2^+) can be obtained. However, the only reaction channel from 2 is an evolution that isomerizes to the more stable intermediate 1 as shown in Figure 4. Not only can the nitrogen atom of HCN⁺ attack C₂H₄, but the carbon atom in HCN⁺ can also attack C₂H₄ through the two attack patterns mentioned above. Unfortunately, we cannot locate the intermediate, which can be seen as formed via the attack of one carbon of C_2H_4 by the carbon of HCN⁺. However, we get a three-membered ring species 3 (NcCHCH₂CH₂⁺), which can be considered to form via the direct attack of the carbon atom of HCN⁺ at the C=C double bond of C_2H_4 . The initial adducts 1 (HCNCH₂CH₂⁺) and 3 (NcCHCH₂CH₂⁺) lie 104.6 and 72.2 kcal/mol, respectively, lower in energy than reactant **R** (HCN⁺ + C_2H_4), which means that the initial association provides the two isomers with enough energy to take subsequent isomerization or dissociation reactions. In the following passage, we mainly discuss the formation pathways of various products formed via 1 and 3.

3.2. Isomerization and Dissociation Pathways. *I. Reaction Product* P_1 (*HCN* + $C_2H_4^+$). The charge-transfer product P_1 (HCN + $C_2H_4^+$) is 70.4 kcal/mol more stable than the reactant. From Figure 4, we find that only one pathway is possible.

Path $\mathbf{P}_1 : \mathbf{R} \to \mathbf{1} \to \mathbf{P}_1$

The initially formed intermediate 1 (HCNCH₂CH₂⁺) can directly dissociate to P_1 (HCN + $C_2H_4^+$) via the cleavage of the N-C single bond. At the B3LYP/6-311G(d,p) level, no transition state is found during this process $(1 \rightarrow P_1)$ in spite of numerous attempts. Alternatively, we obtain the transition state TS1/P₁ at the HF/6-311G(d,p) level, and the structure is shown in Figure 3. We attempt to reoptimize it at the MP2/6-311G(d,p) and QCISD/6-311G(d) levels. However, it does not succeed and often leads to the dissociated product P₁. This again indicates that the dissociation of 1 to P₁ is barrierless, and TS1/P₁ may be an artifact of the HF method.







Figure 3. Part 3 of 3. Optimized structures of the transition states. Distances are given in angstroms and angles are given in degrees.

II. Reaction Product P_2 (*HCNCHCH*₂⁺ + *H*). P_2 (HC-NCHCH₂⁺ + H) lies 70.1 kcal/mol below the reactant. From Figure 4, we find that five formation pathways are possible.

Path $P_2(1): \mathbb{R} \to 1 \to P_2$ Path $P_2(2): \mathbb{R} \to 1 \to 6 \to P_2$ Path $P_2(3) : \mathbb{R} \to 1 \to 6 \to 7 \to P_2$ Path $P_2(4) : \mathbb{R} \to 1 \to 4 \to 7 \to P_2$ Path $P_2(5) : \mathbb{R} \to 3 \to 4 \to 7 \to P_2$

1 (HCNCH₂CH₂⁺) can produce P_2 through the elimination of one of the hydrogen atoms bonded to the internal carbon as

TABLE 1: Zero-Point (kcal/mol), Total (a.u.), and Relative Energies in Parentheses (kcal/mol) and Those Including Zero-Point Vibration Energies (ZPVE) of the Reactant and Products for the $HCN^+ + C_2H_4$ Reaction

species	B3LYP	ZPVE	CCSD(T)	CCSD(T) + ZPVE
R $(^{2}\text{HCN}^{+} + {}^{1}\text{C}_{2}\text{H}_{4})$	-171.5692722	41.10053 (0.0)	-171.2177219 (0.0)	0.0
P_1 (HCN + $C_2H_4^+$)	-171.6871249	40.80152(-0.3)	-171.3294402 (-70.1)	-70.4
\mathbf{P}_2 (HCNCHCH ₂ ⁺ + H)	-171.6764678	38.54916 (-2.6)	-171.3254103 (-67.6)	-70.1
\mathbf{P}_3 (NCCH ₂ + CH ₃ ⁺)	-171.6232905	39.0066 (-2.1)	-171.2315663 (-8.7)	-10.8
$P_4 (CN + C_2H_5^+)$	-171.6201409	40.90124 (-0.2)	-171.2730457 (-34.7)	-34.9
\mathbf{P}_5 (NCCHCH ₂ ⁺ + H ₂)	-171.670772	36.98003 (-4.1)	-171.3066598 (-55.8)	-59.9
\mathbf{P}_{6} (HNCCHCH ₂ ⁺ + H)	-171.695308	38.57216 (-2.5)	-171.3422801 (-78.2)	-80.7
$\mathbf{P}_7 (\text{c-CHCCH}_2\text{N}^+ + \text{H}_2)$	-171.6057926	37.37759 (-3.7)	-171.2587270 (-25.7)	-29.5
\mathbf{P}_{8} (c-NHCCH ₂ C ⁺ + H ₂)	-171.590688	37.49424 (-3.6)	-171.2417497 (-15.1)	-18.7
$\mathbf{P}_{9} (\mathrm{HNCCCH}^{+} + \mathrm{H}_{2} + \mathrm{H})$	-171.6020162	29.76482 (-11.3)	-171.2466805 (-18.2)	-29.5

TABLE 2: Zero-Point (kcal/mol), Total (a.u.), and Relative Energies in Parentheses (kcal/mol) and Those Including Zero-Point Vibration Energies (ZPVE) of the Intermediates and Transition States for the $HCN^+ + C_2H_4$ Reaction

species	B3LYP	ZPVE	CCSD(T)	CCSD(T) + ZPVE
reactant	-171.5692722	41.10053 (0.0)	-171.2177219(0.0)	0.0
1	-171.7402572	44.08005 (3.0)	-171.3891591 (-107.6)	-104.6
2	-171.7090139	45.63058 (4.5)	-171.355261 (-86.3)	-81.8
3	-171.6920361	44.23658 (3.1)	-171.3377775 (-75.3)	-72.2
4	-171.7266609	46.24328 (5.1)	-171.3749748 (-98.7)	-93.5
5	-171.6962873	43.74696 (2.6)	-171.3345371 (-73.3)	-70.7
6	-171.763167	43.80276 (2.7)	-171.4039506 (-116.9)	-114.2
7	-171.7729872	44.84316 (3.7)	-171.417166 (-125.2)	-121.4
8	-171.7295107	47.17293 (6.1)	-171.3792192 (-101.3)	-95.3
9	-171.7556166	43.86718 (2.8)	-171.402578 (-116.0)	-113.2
10	-171.7793418	44.05545 (3.0)	-171.420203 (-127.1)	-124.1
11	-171.7665286	45.50643 (4.4)	-171.4039864 (-116.9)	-112.5
12	-171.6748028	45.68124 (4.6)	-171.3284639 (-69.5)	-64.9
13	-171.5547009	35.45405 (-5.6)	-171.2049832 (8.0)	2.3
TS1/2	-171.702453	44.27512 (3.2)	-171.3458609 (-80.4)	-77.2
TS1/4	-171.6928729	43.73375 (2.6)	-171.3392289 (-76.2)	-73.6
TS1/5	-171.6529215	42.77802 (1.7)	-171.2982557 (-50.5)	-48.9
TS1/6	-171.6911116	41.21233 (0.1)	-171.3299641 (-70.4)	-70.3
TS1/P ₂	-171.6700632	39.49238 (-1.6)	-171.3155628 (-61.4)	-63.0
TS3/4	-171.686515	44.46459 (3.4)	-171.334576 (-73.3)	-70.0
TS4/7	-171.7002572	44.31786 (3.2)	-171.3425111 (-78.3)	-75.1
TS4/8	-171.6278185	42.68247 (1.6)	-171.2772497 (-37.4)	-35.8
TS4/12	-171.6326134	42.08896 (1.0)	-171.2784816 (-38.1)	-37.1
TS4/P ₇	-171.6013036	39.16047 (-1.9)	-171.246473 (-18.0)	-20.0
TS5/P5	-171.596166	38.13531 (-3.0)	-171.2340771 (-10.3)	-13.2
TS6/7	-171.6744336	41.808 (0.7)	-171.3162892 (-61.9)	-61.1
TS6/P ₂	-171.6767881	38.99963 (-2.1)	-171.3236349 (-66.5)	-68.6
TS7/P ₂	-171.6759573	39.00152 (-2.1)	-171.3209441 (-64.8)	-66.9
TS8/9	-171.6889095	44.71437 (3.6)	-171.3354826 (-73.9)	-70.3
TS8/P 8	-171.5775829	40.12152 (-1.0)	-171.2210284 (-2.1)	-3.1
TS9/10	-171.7133394	41.69848 (0.6)	-171.352792 (-84.8)	-84.2
TS9/P ₆	-171.689733	39.33757 (-1.8)	-171.3335571 (-72.7)	-74.5
TS10/11	-171.6331169	40.19793 (-0.9)	-171.2682485 (-31.7)	-32.6
TS11/P ₆	-171.6887558	39.43989 (-1.7)	-171.3276335 (-69.0)	-70.6
TS13/P9	-171.4995805	32.23249 (-8.9)	-171.1384763 (49.7)	40.9
TSP ₆ /13	-171.5429315	34.54867(-6.6)	-171.1924191 (15.9)	9.3

in path $P_2(1)$. 1 can also isomerize into 6 (HCNCHCH₃⁺) via a 1,2 hydrogen shift. Then 6 either undergoes a hydrogenelimination process that leads to P_2 as in path $P_2(2)$ or continuously isomerizes into 7 (H₂CNCHCH₂⁺) via a 1,4 hydrogen shift followed by a hydrogen-elimination process to generate P_2 as in path $P_2(3)$. In addition, 1 can transform to the four-membered ring intermediate 4 (c-CHCH₂CH₂N⁺) via a ring-closure process as in path $P_2(4)$, whereas, in path $P_2(5)$, 4 can be formed via 3 (N-cCHCH₂CH₂⁺) with a negligible barrier of 2.2 kcal/mol. Subsequently, 4 undergoes a ring-opening process that gives rise to 7 (H₂CNCHCH₂⁺) followed by a hydrogen elimination that leads to P_2 .

In path $P_2(1)$, only one barrier of 41.6 kcal/mol needs to be overcome in order for $1 \rightarrow P_2$ to form P_2 . Yet, in path $P_2(2)$, there are two barriers, which are 34.7 and 45.6 kcal/mol for 1 → 6 and 6 → P₂ conversions, respectively. Then we expect that path P₂(1) should be more competitive than path P₂(2). Considering the high-energy barrier of 54.5kcal/mol for the 7 → P₂ conversion that is involved in the latter three pathways, we can deduce that path P₂(3)–(5) cannot compete with path P₂(1), and path P₂(1) should be the optimal channel to form P₂.

III. Reaction Product P_3 (NCCH₂ + CH₃⁺). 1 (HCN-CH₂CH₂⁺) can isomerize to 5 (NCCH₂CH₃⁺) via a concerted 1,4 hydrogen shift, and through the $-CN \rightarrow -NC$ conversion process, the energy barrier for step $1 \rightarrow 5$ is 55.7 kcal/mol. Despite numerous attempts, we cannot locate the transition state for the $5 \rightarrow P_3$ conversion; moreover, we notice that the terminal $C-C_{Me}$ (note that C_{Me} denotes the methyl carbon) bond of 5 is 1.736 Å and is easy to be broken. So, we expect that this may be a single C–C bond-rupture process. The formation pathway of P_3 can be written as

Path
$$P_3: \mathbb{R} \to 1 \to 5 \to P_3$$

IV. Reaction Product P_4 (*CN* + $C_2H_5^+$). For P_4 (*CN* + $C_2H_5^+$), we find that two pathways are possible:

Path
$$P_4(1): \mathbb{R} \to 1 \to 4 \to 12 \to P_4$$

Path $P_4(2): \mathbb{R} \to 3 \to 4 \to 12 \to P_4$

The formation of the four-membered ring intermediate 4 (c-CHCH₂CH₂N⁺) has been discussed in section II. For brevity, we decide not to discuss it again. Via a hydrogen migration along with the ring-opening process, 4 can transform to 12 (CNCH₂CH₃⁺) with a barrier of 56.4 kcal/mol followed by direct dissociation to P_4 . Because the barrier of 31.0 kcal/mol for 1 \rightarrow 4 conversion is significantly higher than 2.2 kcal/mol for 3 \rightarrow 4 conversion, we can deduce that path $P_4(2)$ is the favorite channel to generate P_4 .

V. Reaction Product P_5 (*NCCHCH*₂ ⁺ + H_2). For product P_5 (NCCHCH₂ ⁺ + H_2), only one feasible pathway is located, which can be depicted as

Path
$$P_5(1): \mathbb{R} \to 1 \to 5 \to P_5$$

The formation of **5** (NCCH₂CH₃⁺) is the same as path **P**₃. Subsequently, **5** can produce **P**₅ via a H₂ molecule elimination process. The barrier heights for the steps of $1 \rightarrow 5$ and $5 \rightarrow P_5$ are 55.7 and 57.5 kcal/mol, respectively.

VI. Reaction Product P_6 (*HNCCHCH*₂⁺ + *H*). There are four feasible pathways to form the lowest-lying product P_6 (HNCCHCH₂⁺ + H), which can be written as follows:

Path
$$P_6(1): \mathbb{R} \to 1 \to 4 \to 8 \to 9 \to P_6$$

Path $P_6(2): \mathbb{R} \to 3 \to 4 \to 8 \to 9 \to P_6$
Path $P_c(3): \mathbb{R} \to 1 \to 4 \to 8 \to 9 \to 10 \to 11 \to P_c$

Path
$$P_6(4): \mathbb{R} \to 3 \to 4 \to 8 \to 9 \to 10 \to 11 \to P_6$$

The formation of 4 (c-CHCH₂CH₂N⁺) has been discussed in section II. 4 can convert to 8 (c-NHCH₂CH₂C⁺) via a hydrogenmigration process followed by a ring-opening process to give rise to 9 (HNCCH₂CH₂⁺). The energy barriers of steps $4 \rightarrow 8$ and $8 \rightarrow 9$ are 57.7 and 25.0 kcal/mol, respectively. Subsequently, 9 can either take a hydrogen-elimination process to form P₆ as in path P₆(1) and path P₆(2) or undergo the successive 1,2- and 1,4 hydrogen-shift interconversion to 10 (HNCCHCH₃⁺) and then to 11 (H₂NCCHCH₂⁺), which can finally take a direct N–H bond cleavage to give product P₆ as in path P₆(4).

The barrier height of $1 \rightarrow 4$ (31.0) is much higher than that of $3 \rightarrow 4$ (2.2); thus, path $P_6(2)$ and path $P_6(4)$ are more competitive than path $P_6(1)$ and path $P_6(3)$, respectively. Moreover, we notice that, for $9 \rightarrow P_6$ conversion, only one barrier of 38.7 kcal/mol ($9 \rightarrow P_6$) must be overcome in path $P_6(2)$. However, in path $P_6(4)$, three barriers for $9 \rightarrow 10$ (39.0), $10 \rightarrow 11$ (91.5), and $9 \rightarrow P_6(41.9)$ must be overcome. So, we conclude that the optimal channel to generate P_6 is path $P_6(2)$.

VII. Reaction Product P_7 (c-CHCCH₂N⁺ + H₂). P_7 (c-CHCCH₂N⁺ + H₂) is a four-membered ring product, and we find two pathways are possible:

Path
$$P_7(1): \mathbb{R} \to 1 \to 4 \to P_7$$

Path $P_7(2): \mathbb{R} \to 3 \to 4 \to P_7$

The four-membered ring intermediate 4 (c-CHCH₂CH₂N⁺) can produce \mathbf{P}_7 via a H₂-elimination process with a barrier of 73.5 kcal/mol. Because the barrier of $\mathbf{3} \rightarrow \mathbf{4}$ (2.2) is much lower than that of $\mathbf{1} \rightarrow \mathbf{4}$ (31.0), the optimal channel to form \mathbf{P}_7 is path \mathbf{P}_7 (2).

VIII. Reaction Product P_8 (*c*-*NHCCH*₂*C*⁺ + *H*₂). Product P_8 (*c*-HNCCH₂*C*⁺ + *H*₂), another four-membered ring product, is 18.7 kcal/mol more stable than the reactant. From the PES, we find that two pathways are possible.



Figure 4. Sketch map of the potential energy surface (PES).

Path
$$P_8(1): \mathbb{R} \to 1 \to 4 \to 8 \to P_8$$

Path $P_8(2): \mathbb{R} \to 3 \to 4 \to 8 \to P_8$

The formation of **8** (c-NHCH₂CH₂C⁺) is the same as that in path $P_6(1)$. Then **8** undergoes a H₂ elimination that gives rise to P_8 with a high barrier of 92.2 kcal/mol. Comparing the energy barriers of $1 \rightarrow 4$ (31.0) and $3 \rightarrow 4$ (2.2), we can deduce that path $P_8(2)$ is more competitive than path $P_8(1)$.

*IX. Reaction Product P*₉ (*HNCCCH*⁺ + H_2 + H). P₉ is the secondary product of P₆. It can be formed in the following way:

ath
$$\mathbf{P}_9: \mathbf{P}_6 \rightarrow \mathbf{13} \rightarrow \mathbf{P}_9$$

P₆ is HNCCHCH₂⁺ + H, and the way to generate it has been discussed in section VI. **P**₆ can isomerize to **13** (HNCCH₂CH⁺ + H) via a 1,2 hydrogen shift, and then **13** undergoes a H₂-elimination process that leads to **P**₉. The energy barriers for **P**₆ \rightarrow **13** and **13** \rightarrow **P**₉ are 90.0 and 38.6 kcal/mol, respectively.

3.3. Reaction Mechanism. In the preceding sections, we have obtained nine products, which are P_1 (HCN + $C_2H_4^+$), P_2 (HCNCHCH₂⁺ + H), P_3 (NCCH₂ + CH₃⁺), P_4 (CN + $C_2H_5^+$), P_5 (NCCHCH₂⁺ + H₂), P_6 (HNCCHCHC₂⁺ + H), P_7 (c-CHCCH₂N⁺ + H₂), P_8 (c-NHCCH₂C⁺ + H₂), and P_9 (HNC-CCH⁺ + H₂ + H). For easier discussion, the optimal channels for these nine products are listed again:

Path
$$P_1: \mathbb{R} \to 1 \to P_1$$

Path $P_2(1): \mathbb{R} \to 1 \to P_2$
Path $P_3: \mathbb{R} \to 1 \to 5 \to P_3$
Path $P_4(2): \mathbb{R} \to 3 \to 4 \to 12 \to P_4$
Path $P_5: \mathbb{R} \to 1 \to 5 \to P_5$
Path $P_6(2): \mathbb{R} \to 3 \to 4 \to 8 \to 9 \to P_6$
Path $P_7(2): \mathbb{R} \to 3 \to 4 \to 8 \to P_8$
Path $P_8(2): \mathbb{R} \to 3 \to 4 \to 8 \to P_8$
Path $P_9: P_6 \to 13 \to P_9$

After comparing every optimal channel for forming possible products, we can deduce that the order of energy barriers of the rate-controlling steps increases as follows: path $P_1(1)$ (0.0) \rightarrow path $P_2(1)$ (41.6) \rightarrow path $P_3(55.7) \rightarrow$ path $P_4(2)$ (56.4) \rightarrow path \mathbf{P} (57.5) \rightarrow path $\mathbf{P}_6(2)$ (57.7) \rightarrow path $\mathbf{P}_7(2)$ (73.5) \rightarrow path P_9 (90.0) \rightarrow path $P_8(2)$ (92.2). The much higher barriers make path $P_7(2)$, path $P_8(2)$, and path P_9 kinetically unfeasible. As for the remaining six channels, path P_1 is kinetically the most favorable channel, followed by path $P_2(1)$, yet it is much less competitive compared to path P_1 . Paths P_{3-6} are kinetically the least feasible channels. Because the energy barriers of paths P_{3-6} are very close, within 2 kcal/mol, these four pathways may have contributions comparable to that of the HCN^+ + C_2H_4 reaction. The actual yields of these four products may depend on the reaction conditions in the experiment. As a result, reflected in the final product distributions, we predict that a total of six kinds of products, \mathbf{P}_1 (HCN + $\mathbf{C}_2\mathbf{H}_4^+$), \mathbf{P}_2 (HCNCHCH₂⁺ + H), P_3 (NCCH₂ + CH₃⁺), P_4 (CN + C₂H₅⁺), P_5 (NCCHCH₂⁺ + H₂), and P₆ (HNCCHCH₂⁺ + H), should be observed. P₁ may be the most feasible product with the largest yields. P_2 may be the second most favorable product but it is less competitive than P₁, P₃, P₄, P₅, and P₆, which have the lowest yields observed.

4. Comparison with Experiments

In 2004, Anicich et al. performed experimental studies on the reaction $HCN^+ + C_2H_4$ using the flowing after glow-selected FA-SIFT at room temperature. The experimental results show the products and distributions are in the order of $C_2H_4^+$ + HCN $(0.73) > H_4C_3N^+ + H(0.13) = C_2H_5^+ + CN(0.13) > HC_3NH^+$ + H₂ + H (0.01). Among these products, C₂H₄⁺ + HCN corresponds to P_1 in our results, and the distribution is highest in all of the products. $H_4C_3N^+ + H$ corresponds to two products in our results, which are P2 and P6. P2 occupies a larger part, and $C_2H_5^+$ + CN and HC_3NH^+ + H_2 + H correspond to P_4 and P₉, respectively. On the whole, this result is consistent with our theoretical result. However, there are some discrepancies, that is, (a) as shown in Figure 4, the relative energies of P_2 (HCNCHCH₂⁺ + H) and P_4 (CN + C₂H₅⁺) are -70.1 and -34.9 kcal/mol, respectively. In addition, energy barriers of the rate-controlling step in the formation pathways of P2 and P4 are 41.6 and 56.4 kcal/mol, respectively. Taking thermodynamics and kinetics into consideration, we believe that $H_4C_3N^+$ + H should have a larger branching ratio than that of $C_2H_5^+$ + CN. The branching ratio of $C_2H_5^+$ + CN obtained by Anicich et al. may be overestimated. $C_2H_5^+$ + CN may come from another source. Notice that trace amounts of water are always present in the reaction system. As a result, H_3O^+ may react with C_2H_4 to give $C_2H_5^+$. On the other hand, dissociation of HCN and all other nitriles may yield CN radical. (b) Our calculations show that products P_3 (NCCH₂ + CH₃⁺) and P_5 (NCCHCH₂⁺ + H₂), which were completely ignored by Anicich et al., may also contribute to the final products. The absence of a measurable signal in these cases cannot be taken as proof of the insignificance of these possible products in the reaction $HCN^+ + C_2H_4$ because the sensitivity of their experimental system is not known for these species. Moreover, the existence of NCCH2 and NCCHCH2⁺ in interstellar space has been suggested by previous studies.^{46,47} Therefore, further experimental investigation is still desirable.

The energies of the reactant, products, intermediates, and transition states may be different at different temperatures. From Figure 4, we find that most of the pathways on the PES are exothermic channels, and an increase in temperature may make it easier to proceed through these channels. However, considering that the low temperature in Titan's atmosphere is about 170 K, and its surface temperature is 96 K,⁴⁸ the energy correction for temperature (from 0 to 170 K) may not play a large role.

5. Conclusions

A detailed doublet PES of the HCN $^+$ + C₂H₄ reaction system is built up at the B3LYP and CCSD(T) (single-point) levels of theory. The main calculated results can be summarized as follows: The reaction initiated by the formation of three minimal isomers 1, 2, and 3 followed by a variety of transformations lead to nine dissociation products. Among the nine products, the charge-transfer product P_1 (C₂H₄⁺ + HCN) is the most favorable product; P_2 (HCNCHCH₂⁺ + H) is the second most feasible product but it is much less competitive than P_1 ; and P_3 $(NCCH_2 + CH_3^+)$, **P**₄ $(CN + C_2H_5^+)$, **P**₅ $(NCCHCH_2^+ + H_2)$, and P_6 (HNCCHCH₂⁺ + H) make up the third most feasible products, yet they are much less competitive than P_1 and P_2 . Other products, namely, P_7 (c-CHCCH₂N⁺ + H₂), P_8 (c-HNCCH₂C⁺ + H₂), and P₉ (HNCCCH⁺ + H₂ + H), are kinetically unfavorable, and their contribution to the final products can be considered negligible. Because the isomers and transition states involved in the most favorable channel of the $HCN^{+} + C_2H_4$ reaction are lower in energy than the reactant, the total reaction is expected to be rapid, as is confirmed by experiment. Our theoretical results are consistent with the experimental investigations, and we hope the results may provide

some useful information for understanding the reaction mechanism of HCN^+ with unsaturated hydrocarbon.

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References and Notes

(1) Sagan, C.; Thompson, W. R. Icarus 1984, 59, 133.

(2) Cousteins, A.; Encrenaz, T.; Bézard, B. Icarus 1993, 102, 240.

(3) Keller, C. N.; Cravens, T. E.; Gan, L. J. Geophys. Res. 1992, 97, 12117.

(4) Lara, L. M.; Lellouch, E.; Loppez-Moreno, J. J.; Rodrigo, R. J. Geophys. Res., [Atmos.] 1996, 97, 12261.

(5) Toublanc, D.; Parisot, J. P.; Brillet, J.; Gautier, D.; Raulin, F.; Mckay, C. P. *Icarus* **1995**, *113*, 2.

(6) Yung, Y. L.; Allen, M.; Pinto, J. P. Astrophys. J. 1984, 55, 465.
(7) Bauer, S. J. Adv. Space Res. 1987, 7, 65.

(8) Cravens, T. E.; Robertson, I. P.; Waite, J. H.; Yelle, R. V.; Kasprzak, W. T.; Keller, C. N.; Ledvina, S. A.; Niemann, H. B.; Luhmann, J. G.; McNutt, R. L.; Ip, W. H.; DeLaHaye, V.; Mueller-Wodarg, I.; Wahlund, J. E.; Anicich, V. G.; Vuitton, V. *Geophys. Res. Lett.* **2006**, *33* (7), Art. No. L07105.

(9) Fulchignoni, M.; Ferri, F.; Angrilli, F.; Ball, A. J.; Bar-Nun, A.; Barucci, M. A.; Bettanini, C.; Bianchini, G.; Borucki, W.; Colombatti, G.; Coradini, M.; Coustenis, A.; Debei, S.; Falkner, P.; Fanti, G.; Flamini, E.; Gaborit, V.; Grard, R.; Hamelin, M. M.; Harri, A.; Hathi, B.; Jernej, I.; Leese, M. R.; Lehto, A.; Lion Stoppato, P. F.; López-Moreno, J. J.; Mäkinen, T.; McDonnell, J. A. M.; McKay, C. P.; Molina-Cuberos, G.; Neubauer, F. M.; Pirronello, V.; Rodrigo, R.; Saggin, B.; Schwingenschuh, K.; Seiff, A.; Simões, F.; Svedhem, H.; Tokano, T.; Towner, M. C.; Trautner, R.; Withers, P.; Zarnecki, J. C. *Nature* **2006**, *438*, 785.

(10) Broadfoot, A.; Sandel, B.; Shemansky, D.; Holberg, J.; Smith, G.; Strobel, D.; McConnell, J.; Kumar, S.; Hunten, D.; Atreya, S.; Donahue, T.; Moos, H.; Bertaux, J.; Blamont, J.; Pomphrey, R.; Linick, S. *Science* **1981**, *212*, 206.

(11) Hanel, R.; Conrath, B.; Flasar, F.; Kunde, V.; Maguire, W.; Pearl, J.; Pirraglia, J.; Samuelson, R.; Herath, L.; Allison, M.; Cruikshank, D.; Gautier, D.; Gierasch, P.; Horn, L.; Koppany, R.; Ponnamperuma, C. *Science* **1981**, *212*, 192.

(12) Samuelson, R.; Nath, N.; Borysow, A. Planet. Space Sci. 1997, 45, 959.

(13) Cravens, T. E.; Keller, C. N.; Gan, L. Proceedings of the Symposium on Titan; ESA: 1992; SP-338, 273.

(14) Banaskiewicz, M.; Lara, M. L.; Rodrigo, R.; López-Moreno, J. J.; Molina-Cuberos, G. J. *Icarus* 2000, 147, 386.

(15) Wahlund, J. E.; Boström, R.; Gustafsson, G.; Gurnett, D. A.; Kurth, W. S.; Pedersen, A.; Averkamp, T. F.; Hospodarsky, G. B.; Personn, A. M.;

Canu, P.; Neubauer, F. M.; Dougherty, M. K.; Eriksson, A. I.; Morooka, M. W.; Gill, R.; André, M.; Eliasson, L.; Müller-Wodarg, I. *Science* **2005**, *308*, 986.

(16) Lilensten, J.; Simon, C.; Witasse, O.; Dutuit, O.; Thissen, R.; Alcaraz, C. *Icarus* **2005**, *174*, 285.

(17) Anicich, V. G.; Milligan, D. B.; Fairley, D. A.; McEwan, M. J. Icarus 2000, 146, 118.

(18) Anicich, V. G.; Wilson, P. F.; McEwan, M. J. J. Am. Soc. Mass Spectrom. 2003, 14, 900. (19) Anicich, V. G.; Wilson, P. F.; McEwan, M. J. J. Am. Soc. Mass Spectrom. 2004, 15, 1148.

(20) Anicich, V. G.; McEwan, M. J. J. Am. Soc. Mass Spectrom. 2006, 17, 544.

(21) Adames, N.; Smith, D. J. Am. Soc. Mass Spectrom. 1976, 21, 349.
(22) Adames, N.; Smith, D. J. Chem. Phys. 1980, 72, 288.

(23) Dobrijevic, M.; Ollivier, J. L.; Billebaud, F.; Brillet, J.; Parisot, J. Acta Astronaut. 2003, 398, 335.

(24) Smith, D.; Adames, N.; Miller, T. J. Chem. Phys. 1978, 69, 308.
(25) Toublanc, D.; Parisot, J.; Briller, J.; Gautier, D.; Raulin, F.; Mckay, C. Icarus 1995, 113, 2.

(26) Alcaraz, C.; Nicolas, C.; Thissen, R.; Zabka, J.; Dutuit, O. J. Phys. Chem. A 2004, 108, 9998.

(27) Sekine, Y.; Imanaka, H.; Matsui, T.; Khare, B. N.; Bakes, E. L. O.; Mckay, C. P.; Sugita, S. J. *Icarus* **2008**, *194*, 186.

(28) Redondo, P.; Pauzat, F.; Ellinger, Y. Planet. Space Sci. 2006, 54, 181.

(29) Marco, D. S.; Marzio, R.; Antonio, S. *Chem. Phys.* 2004, 297, 121.
(30) Alves, T. V.; de Oliveria, A. G. S.; Ornellas, F. S. *Chem. Phys. Lett.* 2008, 457, 36.

(31) Milligan, D. B.; Freeman, C. G.; Maclagan, R. G. A. R.; McEwan, M. J.; Wilson, P. F.; Anicich, V. G. J. Am. Soc. Mass Spectrom. 2001, 12, 557.

(32) Osamura, Y.; Petrie, S. J. Phys. Chem. A 2004, 108, 3615.

(33) Anicich, V. G.; Wilson, P. F.; McEwan, M. J. J. Am. Soc. Mass Spectrom. 2004, 15, 1148.

(34) Li, Y.; Liu, H. L.; Huang, X. R.; Geng, C. Y.; Sun, C. C.; Tang, A. C. J. Phys. Chem. A **2008**, *112*, 2693.

(35) Li, Y.; Liu, H. L.; Huang, X. R.; Wang, D. Q.; Sun, C. C.; Tang, A. C. J. Phys. Chem. A **2008**, 112, 8188.

(36) Frisch, M. J.; Trucks, G. W.; Schlegel, H. B.; et al. *Gaussian 98*, Revision A.6; Gaussian, Inc.: Pittsburgh, PA, 1998.

(37) Becke, A. D. J. Chem. Phys. 1993, 98, 5648.

(38) McLean, A. D.; Chandler, G. S. J. Chem. Phys. 1980, 72, 5639.
(39) Krishnan, R.; Binkley, J. S.; Seeger, R.; Pople, J. A. J. Chem. Phys. 1980, 72, 650.

(40) Frisch, M. J.; Pople, J. A.; Binkley, J. S. J. Chem. Phys. 1984, 80, 3265.

(41) Pople, J. A.; Head-Gordon, M.; Raghavachari, K. J. Chem. Phys. 1987, 87, 5968.

(42) (a) Gonzalez, C.; Schlegel, H. B. J. Chem. Phys. 1989, 90, 2154.
(b) Gonzalez, C.; Schlegel, H. B. J. Chem. Phys. 1990, 94, 5523.

(43) (a) The performance of the B3LYP functional for geometry optimization was assessed, for example, by the authors of several versatile model chemistries. G2(B3LYP/MP2/CC): Bauschlicher, C. W.; Partridge, H. J. Chem. Phys. **1995**, 103, 1788. (b) G2M: Mebel, A. M.; Morokuma, K.; Lin, M. C. J. Chem. Phys. **1995**, 103, 7414. (c) G3//B3LYP: Baboul, A. G.; Curtiss, L. A.; Redfern, P. C.; Raghavachari, K. J. Chem. Phys. **1999**, 110, 7650. (d) G3X: Curtiss, L. A.; Redfern, P. C.; Raghavachari, K.; Pople, J. A. J. Chem. Phys. **2001**, 114, 108.

(44) Tokmakov, I. V.; Lin, M. C. J. Phys. Chem. A 2002, 106, 11309.
(45) Xie, H. B.; Ding, Y. H.; Sun, C. C. J. Theor. Comput. Chem. 2005, 4, 1029.

(46) Nesbitt, F. L.; Marston, G.; Stief, L. J. First International Conference on Laboratory Research for Planetary Atmospheres; pp 244–248.

(47) Manaszkiewicz, M.; Lara, L. M.; Rodrigo, R.; López-Moreno, J. J.; Molina-Cuberos, G. J. *Icarus* **2000**, *147*, 386.

(48) Mousis, O.; Gautier, D.; Coustenis, A. Icarus 2002, 159, 156.

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