## Internal Conversion Coefficients in the Decay of Au<sup>199</sup>

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(Received 19 November 1962; revised manuscript received 26 December 1962)

The intensities of the  $\gamma$  rays and electron shell vacancies in the decay of Au<sup>199</sup> have been measured by differential  $\gamma$ -ray spectrometry and  $4\pi \beta \gamma$  coincidence counting. These intensities, per Au<sup>199</sup> disintegration, are 0.368 $\pm$ 0.011, 0.0774 $\pm$ 0.0022, and 0.0031 $\pm$ 0.0015 for the 0.158-, 0.209-, and 0.051-MeV  $\gamma$  rays, respectively, and  $0.208\pm0.014$  and  $0.176\pm0.005$  for the *L*- and *K*-shell vacancies, respectively. Internal conversion coefficients, obtained from these data and the relative conversion electron yields of Nail, Baird, and Haynes, are: in the *K* shell,  $0.317 \pm 0.013$  and  $0.768 \pm 0.032$  for the 0.158- and 0.209-MeV  $\gamma$  rays, respectively; and in the L shell, 0.464 $\pm$ 0.039, 0.139 $\pm$ 0.012, and 8.5 $\pm$ 4.1 for the 0.158-, 0.209-, and 0.051-MeV  $\gamma$ rays, respectively. These values are consistent with the internal conversion coefficients calculated by Rose. The intensities computed from the data for the 0.251-, 0.302-, and 0.460-MeV  $\beta$ -particle transitions are  $0.189 \pm 0.005$ ,  $0.680 \pm 0.025$ , and  $0.131 \pm 0.025$ , respectively.

## **INTRODUCTION**

THE decay of 3.15-day Au<sup>199</sup> has been studied ex-<br>tensively by several workers.<sup>1-5</sup> In all of these<br>investigations, only relative values of the electron and HE decay of 3.15-day Au<sup>199</sup> has been studied extensively by several workers.1-5 In all of these  $\gamma$ -ray intensities were determined and used in the elucidation of the decay scheme, which is shown in Fig. I.<sup>6</sup> In particular, Nail, Baird, and Haynes (to be referred to as NBH) in a recent paper<sup>4</sup> presented a set of average values of the relative conversion electron intensities obtained from a comparison of their data with the earlier results of Cressman and Wilkinson<sup>2</sup> and of Bäckström *et al?* It was then necessary for NBH to use theoretical internal conversion coefficients7,8 with their conversion



<sup>\*</sup> Operated for the Atomic Energy Commission by Union Carbide Nuclear Company.

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- 6 R. W. Bauer, L. Grodzins, and H. H. Wilson, Phys. Rev. 128, 694 (1962). 6  *Nuclear Data Sheets,* compiled by K. Way *et al.* (Printing and
- Publishing Office, National Academy of Sciences—National Re-search Council, Washington, D. C, 1962), NRC 58-9-98, 99.

7 M. E. Rose, *Internal Conversion Coefficients* (North-Holland

electron data in order to compute relative values of the  $\gamma$ -ray transitions in Au<sup>199</sup>.

In the present investigation, the intensities of the  $\gamma$ rays and of the x rays arising from the internal conversion process in Au<sup>199</sup> were measured on an absolute basis. By combining these data with the relative conversion electron intensities of NBH, one can then determine the internal conversion coefficients without reference to any theoretical compilation. It is then also possible to evaluate the total transition intensities in the decay of  $Au^{199}$ .

## **EXPERIMENTAL PROCEDURES**

The Au<sup>199</sup> activity was obtained from the decay of 30min Pt<sup>199</sup>, produced via the Pt<sup>198</sup> $(n, \gamma)$  reaction. Natural Pt was used as the target material since no radioactive isotopes other than Au<sup>199</sup> were formed by the decay of the Pt nuclides produced in the irradiation. The Pt target was dissolved in aqua regia, and the chloride concentration in the solution then adjusted to 8.V. The Au<sup>199</sup> was obtained carrier free by repeatedly extracting it from the Pt solution with ethyl acetate.<sup>9</sup> The ethyl acetate solution was washed with  $6N$  HCl several times. It was then heated to dryness, and the Au<sup>199</sup> residue taken up in *3N* HC1.

Essentially weightless samples were prepared by evaporating known aliquots of the Au<sup>199</sup> solution on thin VYNS film backings; the disintegration rates of these samples were determined by the method of  $4\pi\beta-\gamma$ coincidence counting.<sup>10</sup> These samples were also assayed by differential gamma-ray spectrometry with a  $3$  in. $\times 3$ in. Nal(Tl) detector so that the intensities per disintegration of the *K* x ray, 0.051-, 0.158-, and 0.209-MeV  $\gamma$  rays could be determined.<sup>11</sup> The value of the fluorescence yield,  $\omega_K$ , used for the Hg K x ray was

<sup>&</sup>lt;sup>1</sup> P. M. Sherk and R. D. Hill, Phys. Rev. 83, 1097 (1951). References to earlier work on Au<sup>199</sup> are found in this paper.

<sup>2</sup> P. J. Cressman and R. G. Wilkinson, Phys. Rev. **109,** 872 (1958).

Publishing Company, Amsterdam, 1958). <sup>8</sup>L. A. Sliv and I. M. Band, Leningrad Physico-Technical Institute Report, 1956 [translation: Reports 57ICCK1 (1957) and 58ICCL1, 1958, issued by Physics Department, University of Illinois, Urbana, Illinois (unpublished)].

<sup>9</sup> J. F. Emery and G. W. Leddicotte, *The Radiochemistry o\* Gold*  (National Academy of Sciences—National Research Council), NAS-NS **3036** (1961).

<sup>10</sup> P. J. Campion, Intern. J. Appl. Radiation and Isotopes 4, 232 (1959).

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FIG. 2. The  $\gamma$ -ray spectrum of Au<sup>199</sup>, measured with a 3 in.  $\times$  3 in. NaI(Tl) detector, showing the Gaussian distributions fitted to the photopeaks. The 0.412-MeV  $\gamma$  ray from impurity Au<sup>198</sup> is also shown.

 $0.952 \pm 0.003$ <sup>4</sup> The escape peak correction to the K x ray was taken to be 0.92, $^{\rm 12}$  while the escape peak losses for the 0.158- and 0.209-MeV gammas were less than 0.01. The L-x-ray yield was similarly determined by assaying the samples with a  $1.5$ -in.-diam  $NaI(Tl)$ crystal, 3.4 mm thick; the Be window on the crystal was  $0.137$  g/cm<sup>2</sup> thick. The average fluorescence yield for the *L* shell,  $\tilde{\omega}_L$ , was taken to be 0.391 $\pm$ 0.027; this value is the mean of the values reported by NBH,  $0.410 \pm 0.04$ , and by Haynes and Achor,<sup>13</sup>  $0.371 \pm 0.035$ .

To determine the counting rates of the x and  $\gamma$  rays of Au<sup>199</sup> , the various photopeaks were fitted to Gaussian distributions, after corrections for the Compton distributions of the higher energy gammas were made. To facilitate this process, spectra of monoenergetic gamma rays from divers radioactive isotopes were obtained. Thus, with appropriate changes in high voltage and amplifier gain in the detector system,  $Cr^{51}$  and  $Hg^{203}$ were used to record a spectrum for the 0.209-MeV  $\gamma$  ray: Ce<sup>141</sup> and Sc<sup>47</sup>, for the 0.158-MeV  $\gamma$  ray; Hg<sup>197</sup>, for the K x ray; and the x ray of Sr<sup>85</sup>, for the Au<sup>199</sup> L x ray. Comparison of the resulting net  $K \times \text{ray } (0.072 \text{ MeV})$ and escape peak (0.044 MeV) of Au<sup>199</sup> with those of Hg<sup>197</sup> yielded a very small and approximate contribution from the Au<sup>199</sup> 0.051-MeV  $\gamma$  ray.

The Au<sup>199</sup> spectra, with the resolved Gaussian photopeaks, are shown in Figs. 2 and 3. In order to ascertain that the  $L$ - and  $K$ -x-ray yields were solely those of Au<sup>199</sup>, and not due to Au<sup>198</sup> produced from possible Au<sup>197</sup> contamination of the Pt target, the 0.412-MeV gamma of Au<sup>198</sup> was sought. As shown in Fig. 2, the intensity of this  $\gamma$  ray was very low; the Au<sup>198</sup>/Au<sup>199</sup> ratio was found to be  $0.3\%$ .

## RESULTS AND DISCUSSION

The observed yields of  $\gamma$  rays and electron shell vacancies, in units of intensity per disintegration of Au<sup>199</sup> , are listed in Table I. These quantities are the average values of data accumulated from six Au<sup>199</sup> samples of different activity levels. The stated errors arise from statistical uncertainties due to the radioactive decay process, and to estimated errors encountered in the determination of the areas under the various photopeaks. In addition, the large uncertainty in the L-shell fluorescence yield contributes to the error listed for the

TABLE I. Intensities per disintegration of gamma rays and electron shell vacancies.

Photon or	Experimental
electron hole	results
$L$ shell	$0.208 \pm 0.014$
$K$ shell	$0.176 \pm 0.005$
0.158-MeV $\gamma$	$0.368 \pm 0.011$
0.209-MeV $\gamma$	$0.0774 + 0.0022$
0.051-MeV $\gamma$	$0.0031 + 0.0015$



FIG. 3. The  $\gamma$ -ray spectrum of Au<sup>199</sup>, measured with a 1.5-in.-diam NaI(Tl) crystal, 3.4 mm thick, showing the Gaussian distribution fitted to the Z-x-ray photopeak.

<sup>&</sup>lt;sup>12</sup> P. R. Bell, in *Beta- and Gamma-Ray Spectroscopy*, edited by Kai Siegbahn (Interscience Publishers, Inc., New York, 1955), Chap. V, p. 155. **135.** K. Haynes and W. T. Achor, J. Phys. Radium 16, 635

<sup>(1955).</sup> 

L-shell vacancy. The ratio of 0.158 to 0.209 gammas, equal to  $4.75\pm0.19$ , agrees with the value of  $4.52\pm0.23$ found by Cressman and Wilkinson.<sup>2</sup>

One may express the experimentally determined electron shell vacancies as

$$
e_i = \sum_j \alpha_{ij} N_j = \sum_j e_{ij},
$$
  
(*i* = *K*, *L*; *j* = 0.051, 0.158, 0.209), (1)

where  $e_i$  is the vacancy per disintegration in the *i*thelectron shell,  $N_j$  is the intensity of the *j*th  $\gamma$  ray per disintegration, and  $\alpha_{ij}$  is the *i*th-shell internal conversion coefficient for the  $j$ th  $\gamma$  ray; the vacancy per disintegration of gamma  $j$  in shell  $i$  is  $e_{ij}$ . If relative values of the *eij* are available, the relation

$$
\alpha_{ik} N_k = \alpha_{ij} N_j e_{ik} / e_{ij} \tag{2}
$$

may be used to write Eq. (1) as

$$
e_i = \alpha_{ij} N_j (1 + \sum_{k \neq j} e_{ik}/e_{ij}), \qquad (3)
$$

where the sum of the relative electron vacancies, or conversion electron yields, referred to the jth  $\gamma$  ray, is taken over the other *k* gammas which are observed. Thus, by letting  $j$  in turn represent each of the  $\gamma$  rays of Au<sup>199</sup>, one may determine the  $\alpha_{ij}$ .

Use of the absolute  $\gamma$ -ray and electron shell intensities of Table I, and the relative conversion electron yields of NBH for the  $K$ - and  $L$ -electron shells, in Eq. (3) has led to the experimental values of the  $\alpha_{ij}$  presented in Table II; no knowledge of theoretical internal conversion coefficients was necessary to obtain these results. For comparison, calculated values of the  $\alpha_{ij}$ , obtained by

TABLE II. Internal conversion coefficients,  $\alpha_{ij}$ .

vacancies	Electron Relative conversion Experimental electron vields <sup>b</sup>	$\alpha$ ii	Calculated $\alpha_{ii}$ <sup>e</sup> E <sub>2</sub>	M1
0.158 K 0.209 K 0.158 L 0.209 L 0.051 L	$1.961 + 0.04$ 1.000 $3.06 + 0.07$ $0.193 + 0.006$ $0.471 + 0.01$	$0.317 + 0.013$ $0.768 + 0.032$ $0.464 + 0.039$ $0.139 + 0.012$ $85 + 41$	0.310 0.154 0.475 0.141	0.843 0.129 7.60

<sup>a</sup>  $\alpha_{ij}$  is the internal conversion coefficient of  $\gamma$  ray  $-j$  in electron shell  $-i$ .<br>b Values are from Nals, Baird, and Haynes; see reference 4.<br>v Values are from Rose; see reference 7. As indicated, the values are<br>g

interpolation from Rose's compilation,<sup>7</sup> are also given in Table II. The experimental quantities, within the stated uncertainties, are in good agreement with the *E2* values for the 0.158-MeV  $\gamma$  ray, and with the *M*1 value for the  $0.051$ -MeV  $\gamma$  ray. These results corroborate the findings of NBH, who concluded that the 0.158-MeV transition is a pure *El* transition, and that the fraction of  $E2$  mixing in the 0.051-MeV transition is  $3.3 \times 10^{-4}$ . For the 0.209-MeV transition, the degree of *El* and *Ml*  mixing may be calculated from the experimental values of the  $\alpha_{ij}$  by solving the simultaneous equations

$$
\alpha_{i\angle 09} = a_i E + b_i M, \quad (i = K, L), \tag{4}
$$

where *a* and *b* are the theoretical *El* and *Ml* conversion coefficients of Rose, and *E* and *M* are the fractional values of the *El* and *Ml* transitions, respectively. The values of *E* and *M*, which should sum to unity, are respectively  $0.18 \pm 0.11$  and  $0.88 \pm 0.04$ . These values are comparable to these found by NBH:  $E=0.113\pm0.01$ and  $M=0.887\pm0.01$ .

Using the previously determined  $\alpha_{ij}N_j$ , one may proceed to calculate the total transition probabilities associated with each  $\gamma$  ray, since

$$
T_j = (1 + \alpha_{Tj}) N_j,
$$

where  $\alpha_{Tj}$  is the sum over the *i*-electron shells of the  $\alpha_{ij}$ . Internal conversion coefficients from NBH for the *M, Ny* and *O* shells were also used in the evaluation of the  $\alpha_{Ti}$ . The resulting total transition intensities per disintegration for the 0.158-, 0.209-, and 0.051-MeV  $\gamma$  rays are 0.716 $\pm$ 0.025, 0.153 $\pm$ 0.003, and 0.036 $\pm$ 0.004, respectively. Referring to the decay scheme of Fig. 1, one sees that the intensities per disingegration of the 0.251-, 0.302-, and 0.460-MeV  $\beta$  particles are, respectively,  $0.189 \pm 0.005$ ,  $0.680 \pm 0.025$ , and  $0.131 \pm 0.025$ . The corresponding transition intensities found by Haynes and Achor<sup>13</sup> by directly observing the three  $\beta$ groups are 0.243, 0.693, and 0.064, respectively. It should be also noted that the ratio of the intensities of the 0.251- and 0.302-MeV betas,  $0.278 \pm 0.013$ , from the present work is in agreement with the values determined by NBH,  $0.280 \pm 0.015$ , and by Bauer *et al.*<sup>5</sup>  $0.30 \pm 0.03$ , while the Haynes and Achor ratio is 0.35.