

is possible for  $(k_0, a) \in X$  (III.4) as is easily seen (and  $n^2 \rightarrow \infty$  could only be possible for  $j \rightarrow \infty$ ). We suppose that those  $(k_0, a)$ , for which (3.15) is fulfilled are also in  $X' \subset X$  (III.2). The set  $X'$  of course depends on  $e$  and  $\beta$ ;  $X'(e, \beta)$ , and we could hope that only those  $(k_0, a)$  fulfilling (15) and corresponding to rest states are in the set  $X'(e\beta)$ .

(5) The (4+1) de Sitter group is the group of motion in the de Sitter spherical world<sup>21</sup> with finite extension in space-like and infinite extension in time-like directions. Its curvature tensor is

$$R_{\mu\nu\lambda\rho} = -\lambda^2(g_{\mu\lambda}g_{\nu\rho} - g_{\mu\rho}g_{\nu\lambda}) \quad (3.16)$$

and Einstein's law reads

$$G_{\nu\lambda} = 3\lambda^2 g_{\nu\lambda}, \quad (3.17)$$

where  $\lambda$  is the parameter introduced in (3.3). The radius of the de Sitter world is<sup>21</sup>  $R = 1/\lambda$ , and if we use for  $\lambda$  the empirical value (1.1a) converting MeV into  $\text{cm}^{-1}$  we obtain for our de Sitter world a radius of  $R \approx 10^{-13}$  cm.<sup>22</sup> Thus we can consider a strongly inter-

acting particle as a de Sitter "spherical" world of  $10^{-13}$  cm with finite space-like and infinite time-like extension,<sup>23</sup> a picture which is not too far from our usual image of a particle, which might indicate that our model is not too far from reality.<sup>24</sup>

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<sup>23</sup> If we had chosen the (3+2) de Sitter group [J. B. Ehrman, Proc. Cambridge Phil. Soc. **53**, 290 (1957)] instead of the (4+1), the strongly interacting particle would have been infinite in space but finite in time, in disagreement with our ideas of a particle, but then we would also have obtained  $m^2 = \lambda^2 \alpha^2 - \lambda^2 j(j+1)$  instead of (3.12), which is in obvious disagreement with experimental facts. It should be remarked that the use of the 3+2 deSitter group in the large [see, e.g., C. Fronsdal, Rev. Mod. Phys. **37**, 221 (1965)] is a completely different idea from the present one. There a single mass point is embedded in a large curved universe and has a discrete spectrum which in the limit  $R \rightarrow \infty$  goes over into the continuum states of the mass point in flat space; it is a change of the kinematical group and has nothing to do with dynamics.

<sup>24</sup> For a discussion and interpretation of de Sitter rotator in which the compact subgroup  $R_4$  is diagonalized see A. O. Barut, in Seminar on High Energy Physics and Elementary Particles, Trieste, 1965 (unpublished).

<sup>21</sup> A. S. Eddington, *The Mathematical Theory of Relativity* (Cambridge University Press, New York, 1963), 10th ed., Chap. V

<sup>22</sup> A geometrical interpretation of the contraction process is difficult, as it does not simply mean  $R \rightarrow \infty$  but  $\alpha/R \rightarrow m_0$ .

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