# THERMAL ANALYSIS-MASS SPECTROMETER COMPUTER SYSTEM AND ITS APPLICATION TO THE EVOLVED GAS ANALYSIS OF GREEN RIVER SHALE AND LUNAR SOIL SAMPLES

**EVERETT K. GIBSON, JR.** 

*Planetary and Earth Sciences Dirision, NASA Manned Spacecrafr Center, Houston, Texas 77058 (L'\_ S. A.)*  **(Received 2 March 1972)** 

## *ABSTR4CT*

A thermal analysis-mass spectrometer system controlied by a computer has been developed and successfully used in the analysis of a wide variety of geochemicai samples. The TA-MS computer system provides a very powerful analytical tool for the determination of volatile species released from samples over a very wide temperature and sample-size range. Use of a small laboratory computer and magnetic-tape storage units permits large quantities of analytical information to be handled and easily retrieved\_ The rapid scan capabilities of the quadrupole mass spectrometer operating under computer control are especially useful for following the rapidly changing composition of the evolved gases released from samples during heating under vacuum conditions\_ Samples of volatile-rich Green River shale and two Iunar soils have been anaivzed to show the utility of the TA-MS computer system.

#### **INTRODUCTION**

Thermal analysis-evoived gas analysis can be applied to a wide variety of analytical problems involving thermal decompositions, phase equilibria, and the identity of released volatiles. The mass spectrometer is one of the most powerful analytical tools available to the chemist for the identification of materials that can be easily volatilized or decomposed. When the mass spectrometer is interfaced with a thermal analysis system, an even more powerful analytical tool evolves\_ In the laboratory, a quadrupole mass spectrometer was interfaced with a vacuum-recording thermoanalyzer. This combination has been described previously<sup>1</sup>. The thermal analysis-mass spectrometer (TA-MS) system generates large volumes of information on the composition of the evolved gases, their relative amounts, and temperaturerelease profiles. To handle these data in a convenient manner in which they can be reduced to a usable form in a short period of time, a small laboratory computer with two magnetic-tape units has been interfaced with the mass spectrometer-thermoanalyzer. The complete system consists of the following: (I) a thermoanalyzer vith

**<sup>\*</sup>Presented at the Twenty-Third Pittsburgh Conference on Analytical Chemistry and Applied jpectroscopy. CIcveIand, Ohio, March, 1972.** 

interfaced quadrupole mass spectrometer, (2) a mass spectrometer-computer interface, (3) a PDP  $8/L$  general-purpose digital computer with a tape controller and two magnetic-tape units for data storage, (4) a digital piotter, and (5) a teletype printer. A schematic of the complete system is shown in Fig. 1. Detailed descriptions of a more compIex computer-operated mass spectrometer system have been given by Reynolds *et al\_'\_* 



Fig. 1. Schematic of gas-analysis system.

# **EXPERIMENTAL**

## *Equipment*

A Finnigan 1015S;'L quadrupole mass spectrometer interfaced with a Mettler vacuum-recording thermoanalyzer was used in this work<sup>1</sup>. The computer system used with the TA-MS system was a specially modified System/150, manufactured by Systems Industries. The System/150 is normally used with a gas chromatograph-mass spectrometer system.

The Finnigan mass spectrometer is of the quadrupole type, and the ionized particles are separated according to their mass-to-charge ratio by sweeping an r.f. and d.c\_ voitage across the quadrupole mass-filter assembly. It consists of four molybdenum rods piaced at the corners of a square. Situated at one end of the mass fiiter is an ionizer (rhenium filament operating at 70 eV) the function of which is to ionize the samples and direct the ions into the quadrupole assembly; at the other end of the quadrupole is an electron-multiplier detector. The ions, which have the correct mass-to-charge ratio and have passed through the quadrupole mass filter, impinge upon the first dynode of the ekctron multiplier. This action causes a current flow, and the analog signal is sent to the interface where it is digitized and then sent into the computer. Each voltage sweep sends a complete mass spectrum to the computer. AI1 the data are stored on magnetic tape.

Ekcause the scanning vottage is varied in a linear manner and because of the mass-separating qualities of the hyperbolic electric fields created within the qua-

## 244

drupole rod assembly, the final mass presentation is linear. The spacing between masses I7 and 18 is, for example, identical to that between masses 500 and 501. This condition makes computerization easy compared to magnetic analyzers. The sweep time of the quadrupole mass spectrometer can be as fast as 0.01 sec, making it an ideal instrument for scanning extremely fast eluding gas phases, such as gases rel : ised from the rupturing of vesicles and gas-rich inclusions.

The interface between the computer and mass spectrometer consists of several components. The function of the 15-bit digital-to-analog converter is to select a number typed in by the operator and convert it to an analog rod voltage. By typing in the first and last mass, the operator sets the mass range that the mass spectrometer will scan. The output analog signal from the mass spectrometer is routed to an integrator that integrates this ion current over a length of time that was initially specified by the operator. The output from the integrator is converted to digital form by a 12-bit ana!og-to-digital converter and is then transmitted to the computer, which normalizes and transfers the data onto magnetic tape. The timing generator is capable of generating sample times of 1 msec to 5 min. Also, it **is** capable of automatically adjusting the integration time as a function of signal strength. This capability provides a nearIy constant signal-to-noise ratio over the entire spectrum. The interface also consists of Iogic and registers necessary to interface the mass spectrometer with the computer.

The computer is a small-scale, general-purpose digital system with a I2-bit word length and 4096 words of core storage. The operator communicates through the teletype with the system in a conversational dialogue (a simple yes, no, or numeric responses to computer-program-originated prompting). The software for the system has been developed by Systems Industries. The *executive* program interprets the operator's inputs from the teletype printer, loads the required programs from the magnetic tape into the computer memory, and transfers control to the particular program. *Diagnostic* routines capable of checking the performance of the interface independent of the mass spectrometer and indicating any malfunctions that may occur are aIso avaiIabIe. In addition, there are *utility* and *maintenance* programs consisting of LIST, COPY, WRITE, DUPLICATE, and DELETE routines.

Normal calibration of the mass spectrometer-computer system requires approximately 40 sec. **The calibration reference samples consist of an air-rare-gas mixture (air, helium, argon, and krypton) for the mass** range **l-100** and an organic compound FC-43 (perfiuorotributylamine) for higher mass ranges (50-750). The calibration samples are introduced into the mass spectrometer, the system automatically calibrates on the reference samples, and a tape file with the calibration data (mass number relative to mass set voltage) is originated.

# *Operational modes*

**Two** separate modes of operation possible with the System/l50 are (1) *data*  logging and (2) *spectrum control*. These two approaches differ only in the way that they acquire data; the analytical process is the same for both.

Using the *data logging* approach to data acquisition, the control operations of the mass selectrometer and thermoanalyzer are designated by the operator. The mass spectrometer operates semi-independently of the computer, scanning the spectrum and measuring ion intensity at all mass points within the range of scan. The computer reduces these raw data and presents the analytical results in a mass spectrum in the form of relative abundance versus mass units (or time) on a scale of O-100.

Spectrum control is an approach to data acquisition made possible by the development of the quadrupole mass spectrometer. Using spectrum control, the computer controls the operation of the mass spectrometer as well as analysis of the results. The operator specifies the analytical data to be obtained, and the computer does the necessary tasks. During the analysis of samples, the operator decides the operating conditions for the thermal analysis  $(e.g.,$  heating rate, sample size, weight sensitivity, maximum desired temperature). By matching the operating parameters of the thermoanalyzer and mass spectrometer, the samples are heated at programed rates (linear or isothermal) and the mass spectrometer, under computer control, measures the ion intensities at those mass points that have been designated Other mass points are not sampled. By computer commands to the mass spectrometer, the gases coming from the sample are analyzed at set periods of time  $(e.g.,)$  any time interval between 1 sec and 5 min). The composition of the gases evolved from the sample can be determined as a function of temperature by correlating spectrum number with the recorded sample temperature\_ For example, if the heating rate were  $6^{\circ}$ C/min and the mass spectrometer sampled the gases every 10 sec, a mass spectrum of the gaseous composition wouid be collected at every I "C interval during the heating period.

The selective scanning technique (sampling only designated masses or mass ranges) can increase the effective scan speed of the mass spectrometer for a given sampling rate. This capability is particularly helpful in operating the mass spectrometer in a repetitive scan mode, such as when it is interfaced with the thermoanalyzer or a gas chromatograph. The total amount of data that has to be recorded is reduced, and, by permitting more scans for a given time period, the resolution of the ion-intensity-versus-time information is also improved.

When spectrum control or computer-controlled data acquisition is used, the following parameters for each analysis can be set by the operator:

(I) Scanning continuous and noncontinuous mass ranges (up to eight separate mass ranges) and acquiring only relevant data, thus increasing the effective scan speed\_

(2) Specifying different integration times for different mass ranges, thus enhancing the sensitivity of the instrument

(3) Specifying sampling rates per a-m-u. to detect mass defects and to determine isotopic ratios accurately.

(4) Specifying thresho!d levels to suppress background noise of the entire system.

(5) Using the automatic adjustment of integration times as a function of signal

**strength in the signal optimization mode, resulting in a nearIy constant signal-to-noise ratio** 

**(6) Determining conditions for the thermal analysis-mass spectrometric analysis, specifying the maximum run time and allowing the instrument to remain**  unattended with the assurance that all desired data will be acquired and that no evolved gases will be missed in the mass ranges specified.

## *Output routines*

**The data ourput mode is used to output information stored on the magnetic tape from the TA-MS computer analysis of sampIes. During the mass spectrometer control and data-acquisition phase, the pIotter gives a real-time gas-release pattern. At the end of the analysis, the computer can also plot a reconstructed gas-release pattern which normalizes on the largest peak or gas-release region. The real-time gas-release pattern is simply a plot of the ion current from the electron multiplier at each recorded spectrum scanned. In addition to the real-time gas-release pattern, the following output routines are avaiIabIe:** 

**(1) Reconstruction of a gas-release pattern that normalizes the largest gasrelease peak or region to 100 and indexes the mass spectra acquired during the run.** 

**(2) Printed or plotted spectra from any mass spectrum cohected.** 

**(3) Piots of individual mass (e.g., mass 18) peaks as a function of temperature, providing information on the temperature region of greatest gas release.** 

**(4) Spectrum plots or printouts with positively identified mass scales and numeric ion intensities.** 

**(5) Spectrum plots or printouts with the background spectra subtracted to eliminate the effect of contaminants or backgrounds.** 

**The TA-MS computer system has been used to identify volatiles released during the heating, their abundances, temperature-release ranges, and sequence of release. By combining the evolved gas information with the thermal analysis data, complete characterization of the decomposition path for a wide variety of materials can be obtained. The analytical technique has been successfully apphed to the analysis of lunar samples, meteorites, minerals, terrestrial rocks, and volcanic ashes in addition to volatile-rich materials such as oil-rich shales, paints, plastics, and polymers. A detailed discussion of the analysis of the volatile-rich Green River shale and the volatile-depleted Apollo 14 and 15 soils will be given to show two applications of the TA-MS computer system.** 

# *Green River shale pyrolysis*

**The Green River shaie is an oil-rich material that contains up to 35%**  volatiles<sup>3</sup>. Detailed discussions of the volatiles and organic compounds found in this **oil-rich shale have been given by Forsmart and by Robinson'. Recently, Smith and Johnson6 have discussed simultaneous DTA-TG-MSA studies of Green River shale components and similar oil-rich sha!es. This discussion of Green River shale**  analysis will be restricted to the application of the TA-MS computer system to the analysis of a 6.0-mg sample of the oii-rich shale. The sample analyzed had been ground to pass through a 100-mesh sieve  $(-147 \mu)$ . The sample holder used was an 8-mmdiameter, flat, platinum, dish-type holder resting directly on the Pt/Pt-10% Rh thermocouple. The sample had been evacuated to  $10^{-6}$  torr; a  $10^{\circ}$ C/min heating rate was used. During analysis, the sample weight, temperature, and composition of the evolved gases were recorded. The reconstructed thermal-analysis curve and gasrelease profile for the sample analyzed are shown in Fig. 2.



Fig. 2. Weight-loss curve and gas-release profile from analysis of a 6.0-mg sample of Green River **shale,** 

Green River shale lost weight in two distinct regions. The two large weight Iosses were exactly matched by the two gas-release peaks. Excellent agreement existed between the weight-loss regions and the evolved-gas peaks (Fig. 2). When heated to approximately  $375^{\circ}$ C, the sample began to lose weight and continued to do so until the temperature reached  $450^{\circ}$ C. The weight loss in this temperature interval was 13.3 weight percent. The gas-release patterns shown in Fig. 2 exhibit a peak for the gases between spectrum numbers 71 and 90. The compositions of these evolved gases (spectrum number 77) are shown in Fig. 3. Mass spectrum number 77 shows the typical hydrocarbons of masses 27, 29, 39, 41, 55, 57, and 69. In addition, smaI1 traces of  $H_2O$  (mass 18) and CO and/or  $N_2$  (mass 28) are also present. To determine the temperature range in which the hydrocarbons are released, the release profile for mass 55 (which is a typical hydrocarbon fragment) has been pIotted in Fig. 4. The



Fig. 3. Mass spectrum of gases released at 400°C. Spectrum shows the presence of numerous hydrocarbon fragments along with trace quantities of H<sub>2</sub>O and CO.

GREEN RIVER SHALE 1FSS 10 DEGIC MIN : 4-23 71 60 MG.



Fig. 4. Reconstructed gas-release profile of mass 55. The major peak results from hydrocarbon fragments released in the 390–450 °C temperature range. The minor peak at spectrum numbers 120– **135 results from the pyroIysis of trace quantities of condensed hydrocarbons in the analysis system.** 





Fig. 5. Mass spectrum of gases released at 600 °C. Spectrum shows the presence of only CO<sub>2</sub> resulting **from decomposition of carbonate phases.** 



**Fig. 6. Reconstructed gas-rcicasc profde of mass 44. The gas-release profde shows that CO\* is**  released only between spectrum numbers 100 and 120.

release of the mass 55 is restricted to spectra 71 to 90 with perhaps minor amounts (less than 15% amplitude) released at temperatures above 700°C.

The Green River shale sample loses 24.6 weight percent between 590 and 65O"C\_ This weight loss results from the decomposition of carbonate phases found in the oil-rich shale. The carbonate phases have been discussed by Smith and Johnson<sup>6</sup> and consist of nahcolite, dolomite, dawsonite, and "ferroan". The mass spectra of the evolved gases between temperatures 590 and 650 "C (spectra 100 to 120) are shown in Fig. 5. Mass 44 (CO,) was the only peak that appeared in the spectra. A reconstructed gas-release profile for mass 44 is shown in Fig. 6. The spectrum number given in Fig. 6 is directly correlated with the temperature given in Fig. 2. The single peak for mass 44 between spectrum numbers 100 and 120 indicates that  $CO<sub>2</sub>$  is released only between 590 and 650°C. It is interesting to note that the two small spikes on the top of the  $CO_2$ -release profile (Fig. 6) are very similar to those observed previously for the decomposition of dolomites found in meteorites<sup>1</sup>.

# *Lunar soil pyrolysis*

Lunar samples present a unique requirement that the analyst obtain the maximum amount of useful information on a minimum sample size. The lunar samples from the six previous lunar missions (Apollo I I, 12, 14,15, and Luna 16 and 20) are all depleted in their volatile elements (e.g., C, N<sub>2</sub>, S, Na, and K) and compounds such as hydrates, carbonates and sulfates<sup> $7-10$ </sup>. The lunar soils are complex



Fig. 7. Weight-loss curve and gas-release profile for Apollo 14 soil 14163,178. The presence of H<sub>2</sub>O **and CO2** being released from **the sampIe at** temperatures **below 15O'C shouId be noted. The arrow on the weight-loss curve around 114O'C indicates the initial melting of the sample.** 



Fig. 8. Weight-loss curve and gas-release p.ofile for Apollo 15 soil 15601,31 from Hadley Rille. The spikes on the gas-release curves for CO and N<sub>2</sub> should be noted. These spikes undoubtedly result from rupturing of vesicles and gas-rich inclusions. The release of  $O<sub>z</sub>$  at 1300<sup>°</sup>C results from the dissolution of Fe into the platinum crucible, releasing O<sub>2</sub>. The absence of the sudden release of O<sub>2</sub> **a-hen the san.pIes are heated in an afumina crucible (Fig. 7) should be noted.** 

mixtures of individual mineral grains, glass particles, rock chips, and trace quantities of meteor.tic components. Thermal analysis-inorganic gas-release studies have been performed on samples from each of the four Apollo lunar missions<sup>11.12</sup>; DTA has shown<sup>12</sup> that the lunar soils undergo melting in the region of 1080–1250<sup>°</sup>C.

Because of the depletions in the volatile components in lunar samples, the small weight-loss measurements (TG) on lunar samples are not surprising. A TG and gas-release profile for the Apollo 14 soil collected near the lunar module and for the Apollo 15 soil from Hadley Rille are given in Figs. 7 and 8, respectively. Lunar soils analyzed from the first four Apollo missions lose less than  $0.35 \pm 0.10$  weight percent when heated to 1000°C under vacuum of  $10^{-6}$  torr, and even after 1400°C the soils lose less than  $1.70 \pm 0.50$  weight percent. These weight-loss studies were performed on soil samples ranging in weight from 200 to 225 mg. The soil samples were analyzed in 16-mm-diameter platinum and alumina crucibles. The small weight losses measured are indicative of the past thermal histories of these complex samples.

A separate sample of Apollo 14 soil 14163,178 heated to 1400°C in air gained 4.80 weight percent. The weight gain resulted from the oxidation of iron and other reduced phases in the soil samples. The lunar samples had been formed under extremely low partial pressures of oxygen, and the absence of iron in the ferric state is to be expected<sup> $7-10$ </sup>. The weight gain from heating lunar soil samples in air perhaps couId be used 3s 3n indicator of the relative degrees of reduction for the various **soil** 

252

samples. Hanneman<sup>13</sup> previously reported  $\lambda$  net weight gain when the lunar soils were heated under oxidizing conditions.

The small weight losses that result during heating of lunar soils in vacuum do not encourage one in his attempt to locate and determine the volatile phases released from the lunar soils; however, the high sensitivity of the mass spectrometer shows that numerous volatile gas species are released during the heating of the samples (Figs. 7 and 8). The gas-release patterns obtained from the lunar soils indicate that gaseous species originate from several sources. Evidence has been obtained indicating that the gases are: (I) contaminants, (2) solar-wind-derived species, (3) reaction products, and (4) gas phases within vesicles and gas-rich inclusions<sup>11,12</sup>.

Thermal analysis-inorganic gas-reIease studies of the lunar soils have shown that very little voiatiie-rich materials or hydrous mineral phases are found in the lunar soils<sup> $11,12$ </sup>. The gases released during the heating of the soils result from separate and distinct sources. Water vapor and  $CO<sub>2</sub>$  have been adsorbed by the lunar soils that have been exposed to the terrestrial atmosphere after return to earth. These two gaseous species are loosely bound to the surfaces of the soiis and are lost from the samples when they are heated under vacuum at temperatures below 150°C. The Apollo 14 soil sample (Fig. 7) had been exposed to the atmosphere for 1 month before analysis, while the Apollo 15 soil (Fig. 8) had been stored under  $N_2$  since return from the moon and was not exposed to the atomsphere until the analysis began. The effects of exposure of the lunar soils to the atmosphere can be seen by the differences in the gaseous species released at temperatures below  $150^{\circ}$ C (Figs. 7 and 8).

Solar-wind components such as  $H_2$  and He and possibly  $CO_2$  are trapped within the outer surfaces of individual mineral and glass grains, and these gases are released between 300 and 7OO'C. Additional low-moiccular-weight species such as  $H<sub>2</sub>O$  (possibly in the hydroxyl form) can be formed on the quter surfaces of rocks and mineral grains as the result of proton irradiation from **the** solar wind\_ initial laboratory experimental results from irradiation with lOO-keV protons on lunar analogs produce trace quantities of  $H_2O$ , and the temperature-release profile along with the concentration of  $H<sub>2</sub>O$  produced is similar to that released from lunar samples<sup>12</sup>.

The release of the gaseous species CO,  $CO<sub>2</sub>$ , H<sub>2</sub>S, and SO<sub>2</sub> at temperatures above 600°C result from the reaction products of C- and S-containing phases found with mineral and glass found in the soils. The S-containing gaseous phases ( $SO<sub>2</sub>$  and H,S) are released between I@00 and 1300°C and are derived from reaction of troilite and other sulfides with the silicate phases at elevated temperatures. Identical gasrelease profiles have been produced using lunar analogs<sup>11</sup>. The S-containing gases have not been plotted in Figs. 7 and 8 to prevent cluttering the figures.

The presence of gas phases in vesicles and gas-rich inclusions along with those gases exsolving from the melt is shown by the gas spikes around  $1140^{\circ}$ C in Figs. 7 and 8. The gases present are CO or  $N_2$  or a combination of the two. These gas phases are released as sudden bursts of gas at temperatures immediately above the melting point ( $1130 \pm 20^{\circ}$ C) of the samples. The identity of the gases found are similar to those that

**should be expected from theoretical considerations of the lunar gaseous environment**  during the time of crystallization of the lunar samples<sup> $14$ </sup>.

#### **CONCLUSIONS**

**The analytical technique of combined thermal analysis-mass spectrometric computer control has been successfully developed and applied to the analysis of a**  wide variety of geochemical samples. Samples analyzed have ranged from oil-rich **shaies, which contain large quantities of volatile components, to Itmar soil samples, which have been depleted in their volatile content\_ The TA-MS computer system provides a very powerful tool for the analysis of volatile components released from samples over a very wide temperature and sampie-size range. The use of the small**  laboratory computer and magnetic-tape units for the storage of the analytical data **shows Iarge quantities of information to be reduced easily. The capabihty of the dataoutput system to provide a gas-release profile of a single mass as a function of tempera**ture is very convenient. The rapid scan capabilities of the quadrupole mass spectio**meter are especiahy useful to foliow the rapidly changing composition of the evolved gases. This technique can be used in other fields of analysis where, upon pyrolysis of samples, the composition of the released gas phases as a function of temperature is desired-**

#### **ACKNOWLEDGMENTS**

The assistance of Vic Borgnis of Finnigan Instruments and Pete Olsen of *Systems* **Industries is acknowledged for their efforts in helping make the TA-MS computer system operational. Their expertise in mass spectrometer-computer interfacing and computer programming ahowed for significant time-saving procedures to be used during this investigation. Suzanne M\_ Johnson and Gary W. Moore assisted in data collection.** 

## **REFERENCES**

- 1 E. K. Gibson, Jr. and S. M. Johnson, Thermochim. Acta, 4 (1972) 49.
- 2 W. E. Reynolds, V. A. Bacon, J. C. Bridges, T. C. Coburn, B. Halpern, J. Lederberg, E. C. Levinthal, E. Steed and R. B. Tucker, Anal. Chem., 42 (1970) 1122.
- 3 M. T. J. Murphy, in G. Eglinton and M. T. J. Murphy (Eds.), *Organic Geochemistry*, Springer-*Verlag. New York, 1969,* **p\_ 74\_**
- 4 J. P. Forsman, ir. I. A. Breger (Ed.), Organic Geochemistry, Macmillan Co., New York, 1963, **p\_ IIs\_**
- 5 W. E. Robinson, in G. Eglinton and M. T. J. Murphy (Eds.), Organic Geochemistry, Springer-**Verlq, Sew York, 1969. p\_ 619.**
- 6 J. W. Smith and D. R. Johnson, in R. F. Schwenker, Jr. and P. D. Garn (Eds.), Thermal Analysis, Vol. 2, Academic Press, New York, 1969, p. 1251.
- 7 **L.SPET, Scicrrcc. I65 (1969) 1211.**
- **% LSPET, Science, 167 (1970) 1325\_**
- **9 LSPFI, Science. IT3 (1971) 68i\_**
- **IO ISPFT.** *Science, I75 (1972)* **363.**
- 11 E. K. Gibson and S. M. Johnson, Proc. 2nd Lunar Sci. Corf., Suppl. 2, Vol. 2, Geochim. Cosmo $chim.$  *Acta*, 1971, p. 1351.
- **12 E. K. Gibson and G. W. Moore,** *Proc. 3rd Lunar Sci. Conf, Suppl.3, Vol.* **2,** *Geochim, Cosmochim. Acta, 1972, p- 2029.*
- 13 R. E. Hanneman, Proc. Apollo 11 Lunar Sci. Conf., Suppl. 1, Vol. 2, Cosmochim. Acta, 1970, p\_ *1207.*
- **14 T. R. Wellman.** *Nature. 225 (197@) 716.*