

# Estimate of the Nuclear Moment of $\text{Ni}^{61}$ from Electron Spin Resonance\*

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The electron spin resonance spectra of nickel and cobalt have been studied in single crystals of MgO. Hyperfine structure was detected from the  $\text{Ni}^{61}$  isotope in the spectrum of  $\text{Ni}^{+2}$  and the hyperfine splitting constant  $A^{61}$  found to be  $(8.3 \pm 0.4) \times 10^{-4} \text{ cm}^{-1}$ . X irradiation of cobalt-containing crystals results in the formation of  $\text{Co}^{+1}$  (isoelectronic with  $\text{Ni}^{+2}$ ). The hyperfine splitting constant  $A^{59}$  is  $(54.0 \pm 0.2) \times 10^{-4} \text{ cm}^{-1}$ . Comparison between  $A^{59}$  and  $A^{61}$  yields a value of  $0.30 \pm 0.02 \text{ nm}$  for the nuclear moment of  $\text{Ni}^{61}$ .

THE nuclear spin of  $\text{Ni}^{61}$  was recently determined by Ludwig *et al.*<sup>1</sup> from a study of the electron spin resonance spectrum of nickel as an impurity in germanium. They used nickel enriched in  $\text{Ni}^{61}$  and observed a clearly resolved hyperfine quartet, indicating a spin  $I^{61} = \frac{3}{2}$ . We have studied the electron spin resonance spectrum of  $\text{Ni}^{+2}$  ions in single crystals of MgO and have detected a weak structure which may be attributed to naturally occurring  $\text{Ni}^{61}$ . So far as we are aware, this is the first report of hyperfine structure from unenriched nickel.

At 77°K the spectrum of  $\text{Ni}^{+2}$  in MgO, as reported previously,<sup>2</sup> consists of a single, isotropic broad line (width  $\sim 50$  gauss) with, at high microwave power, a sharp double quantum line superposed at its center. The width of this central line is sufficiently small to permit observation of hyperfine structure even though there is no chance of doing so for the broad (single quantum) line. Figure 1 is a record of the spectrum, showing two weak lines, equally disposed about the central one and separated from one another by  $23.9 \pm 1.2$  gauss. These lines have the same width as the central one and show a similar dependence of intensity on microwave power, i.e., this varies as the square of the power, rather than linearly. Their intensity, as a fraction of the main line, is found to be  $(0.38 \pm 0.10)\%$ , which may be compared with the expected value of 0.31% (the natural abundance of  $\text{Ni}^{61}$  being 1.25%).

We assume these weak lines to be the outer pair of a hyperfine quartet, the inner pair (shown dotted in Fig. 1) being lost in the central line. If this assumption is correct, it leads to a hyperfine splitting constant  $A^{61} = (8.3 \pm 0.4) \times 10^{-4} \text{ cm}^{-1}$ .

To obtain an estimate of the nuclear moment of  $\text{Ni}^{61}(\mu^{61})$ , this hyperfine splitting may be compared with that from a nucleus of known magnetic moment and ideally the comparison should be made on an ion which is isoelectronic with  $\text{Ni}^{+2}(3d^8)$ . One possible choice is  $\text{Co}^{+1}$  and we have observed the spectrum of this ion in MgO at 77°K, following x irradiation of

samples containing  $\text{Co}^{+2}$ . It consists of eight hyperfine lines ( $I^{59} = \frac{7}{2}$ ) spaced approximately 55 gauss apart and, at large microwave powers, each hyperfine component is crowned by a sharp double quantum absorption as in the case of the  $\text{Ni}^{+2}$  line. A recording of the spectrum is shown in Fig. 2. The sharp line between the third and fourth hyperfine components is the double quantum absorption from  $\text{Ni}^{+2}$ ; other lines present arise from  $\text{Fe}^{+3}$ ,  $\text{Mn}^{+2}$ , and  $\text{Cr}^{+3}$ . It is of interest to note that, at low microwave powers, each  $\text{Co}^{+1}$  line shows a sharp dip at its center as we found earlier<sup>2</sup> for  $\text{Ni}^{+2}$ .

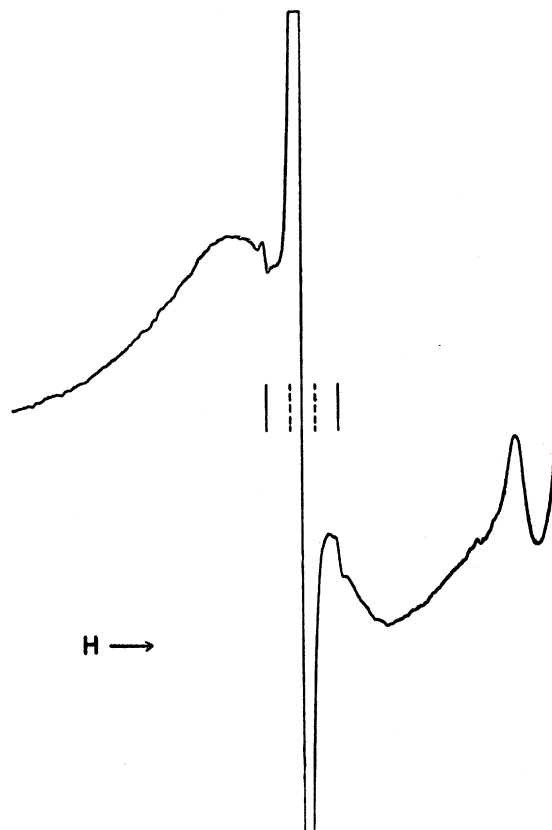


FIG. 1. Spectrum of  $\text{Ni}^{+2}$  in MgO showing a broad line with the sharp, double quantum line superimposed at its center. The position of the  $\text{Ni}^{61}$  hyperfine quartet is indicated at the center of the figure, the dotted pair of lines lying under the central one.

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<sup>1</sup> G. W. Ludwig, H. H. Woodbury, R. O. Carlson, Phys. Rev. Letters 1, 16 (1958).

<sup>2</sup> J. W. Orton, P. Auzins, and J. E. Wertz, Phys. Rev. Letters 4, 128 (1960).

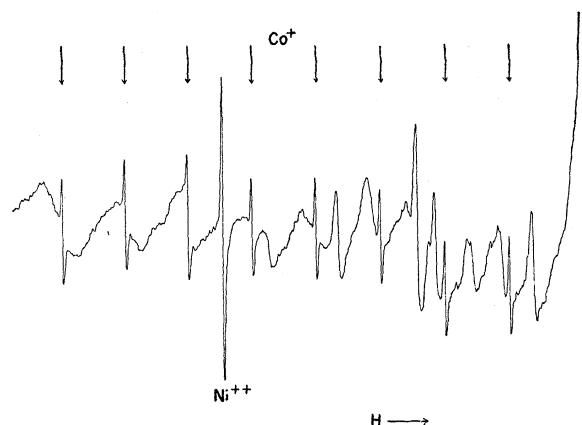


FIG. 2. Spectrum of  $\text{Co}^{+1}$  in  $\text{MgO}$ . The eight sharp lines are double quantum transitions, i.e.,  $\Delta M=2$ ,  $\Delta m=0$ . The double quantum line from  $\text{Ni}^{+2}$  lies between the third and fourth  $\text{Co}^{+1}$  line. Other lines are due to  $\text{Fe}^{+3}$ ,  $\text{Mn}^{+2}$ , and  $\text{Cr}^{+3}$ .

The sharpness of the double quantum lines makes it possible to measure the  $g$  value and hyperfine splitting

with considerable accuracy. The results are:  $g=2.1728 \pm 0.0005$ ,  $A^{59} = (54.0 \pm 0.2) \times 10^{-4} \text{ cm}^{-1}$ .

Using the usual formula  $A = \gamma \beta \beta_N \langle 1/r^3 \rangle_{\text{av}}$ , and writing  $\gamma = \mu/I$ , we calculate the value of  $\mu^{61}$  from

$$\mu^{61} = \frac{A^{61} I^{61}}{A^{59} I^{59}} \mu^{59} R, \quad (1)$$

where  $I^{61} = \frac{3}{2}$ ,  $I^{59} = \frac{7}{2}$  and  $R$  is the ratio  $\langle 1/r^3 \rangle_{\text{av}}(\text{Co}) / \langle 1/r^3 \rangle_{\text{av}}(\text{Ni})$ . From the measured values of  $A$ , we find  $\mu^{61} = 0.31 R \text{ nm}$ .

To estimate a possible value for  $R$  we may calculate it for the case of  $\text{V}^{+2}$  and  $\text{Cr}^{+3}$  in  $\text{MgO}$  using the known values of the nuclear moments of  $\text{V}^{51}$  and  $\text{Cr}^{53}$ . The values of  $A$  are 74.3 and 16.2 ( $\times 10^{-4} \text{ cm}^{-1}$ ), respectively,<sup>3</sup> giving  $R \approx 0.98$ , i.e., very close to unity. Thus, we estimate the nuclear moment of  $\text{Ni}^{61}$  to be  $(0.30 \pm 0.02)$  nuclear magneton. The probable error is somewhat arbitrary.

<sup>3</sup> J. W. Orton, *Reports on Progress in Physics* (The Physical Society, London, 1959), Vol. 22, p. 204.

## Decays of $\text{Rh}^{106}$ and $\text{Ag}^{106}$

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The gamma-ray spectra of 30-sec  $\text{Rh}^{106}$  and 8.3-day  $\text{Ag}^{106}$ , which both decay to  $\text{Pd}^{106}$ , have been studied with scintillation spectrometers. Three gamma-gamma angular correlations have also been measured. The results are consistent with the following level scheme for  $\text{Pd}^{106}$ : 0.513(2+), 1.131(2+), 1.137(0+), 1.360 or 1.213, 1.563(2+), 1.73(2 or 3), 1.84, 1.88, 1.94(3- or 4+), 2.01, 2.052(4+), 2.09(3), 2.28, 2.305(3 or 4), 2.352(4+), 2.46, 2.62, 2.764(5-), 2.87, and 3.08 Mev. The transition between the first and second 2+ levels was found to consist primarily of E2 radiation. The branching ratio obtained for the cascade to crossover gamma rays from the second 2+ level is  $2.1 \pm 0.3$ . This ratio combined with Coulomb excitation data of Stelson and McGowan gives a value of  $1.0 \pm 0.3$  for the ratio  $B(E2, 2' \rightarrow 2)/B(E2, 2 \rightarrow 0)$ .

### I. INTRODUCTION

**T**HIRTY-SECOND  $\text{Rh}^{106}$  decays by beta-ray emission to  $\text{Pd}^{106}$ , and 8.3-day  $\text{Ag}^{106}$  decays by orbital electron capture also to  $\text{Pd}^{106}$ . The spin and parity of  $\text{Rh}^{106}$  have been deduced as 1+ from the comparative half-lives of the beta-ray transitions to the 2+, 0.513-Mev level and 0+ ground state in  $\text{Pd}^{106}$ .<sup>1</sup> The spin of  $\text{Ag}^{106}$  has been measured by Ewbank *et al.*<sup>2</sup> as 6. Alburger and Toppel<sup>3</sup> from their investigation of the de-

cays of  $\text{Rh}^{106}$  and  $\text{Ag}^{106}$  have proposed the energy level diagram given in Fig. 1. Most of the levels are shown as being populated by both  $\text{Rh}^{106}$  and  $\text{Ag}^{106}$ . This is rather surprising in view of the large difference in the spins of the two isotopes. Because of this unsatisfactory situation it was felt desirable to re-examine the decays of  $\text{Rh}^{106}$  and  $\text{Ag}^{106}$ .

The gamma rays of  $\text{Rh}^{106}$  given in Fig. 1 include all previously reported gamma rays<sup>1,4,5</sup> with the exception of a 2.28-Mev gamma ray observed by Kahn and Lyon.<sup>5</sup>

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<sup>1</sup> D. E. Alburger, *Phys. Rev.* **88**, 339 (1952).

<sup>2</sup> W. B. Ewbank, W. A. Nierenberg, H. A. Shugart, and H. B. Silsbee, *Phys. Rev.* **110**, 595 (1958).

<sup>3</sup> D. E. Alburger and B. J. Toppel, *Phys. Rev.* **100**, 1357 (1955).

<sup>4</sup> J. J. Kraushaar and M. Goldhaber, *Phys. Rev.* **89**, 1081 (1953).

<sup>5</sup> B. Kahn and W. S. Lyon, *Phys. Rev.* **92**, 902 (1953).